Spatial Clustering Analysis of the High-Energy Contained Vertex Events



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Each Event Observed Gets Full-sky Reconstruction Treatment

Reconstruction map of a ~60 TeV shower-like event in local coordinates



Reconstruction map of a ~90 TeV track-like event in local coordinates



Directional Reconstruction

- As C. Kopper mentioned in the previous talk, each event gets a full-sky likelihood reconstruction
- Every PMT timing profile used for fit
- Systematic effects are studied and included by further applying a Gaussian smearing
- Cross-check with different fit methods performed

Likelihood Search for a Point Source - Test Statistic (TS) Calculation -

Maximize the likelihood L at every point in the sky x



* Events' energies not used in the likelihood

TS is calculated for every point in the sky x

$$TS(x) = 2 \times \log \left(\frac{L(x)}{L_0}(x) \right)$$

where $L_0 = L(x, n_s = 0)$

Resulting Test Statistic Map



Likelihood Search for a Point Source - Scrambling & Significance Calculation -

- Randomize right ascension (RA) of event reconstruction maps
- Repeat the likelihood calculation at every point in the sky to obtain a TS map
- Compare the highest TS value obtained to the highest TS value actually observed Fraction of times randomized map produced higher TS



- Simulation-free estimation of the significance of the highest TS observed
- Significance derived only from RA clustering (clustering in declination not included in scrambling)
- Significance does NOT include signal/background event separation
- \rightarrow Simply a significance against the hypothesis of uniform event distribution

Conclusions

- First point source analysis in IceCube to include all 3 flavor neutrinos (both track-like and shower-like events)
- First point source analysis in IceCube that pushes the energy threshold in the Southern sky to ~50 TeV
- No evidence of spatial clustering

Stay tuned for a complete overview plenary talk

by N. Whitehorn tomorrow!