

# **BATSE – History**

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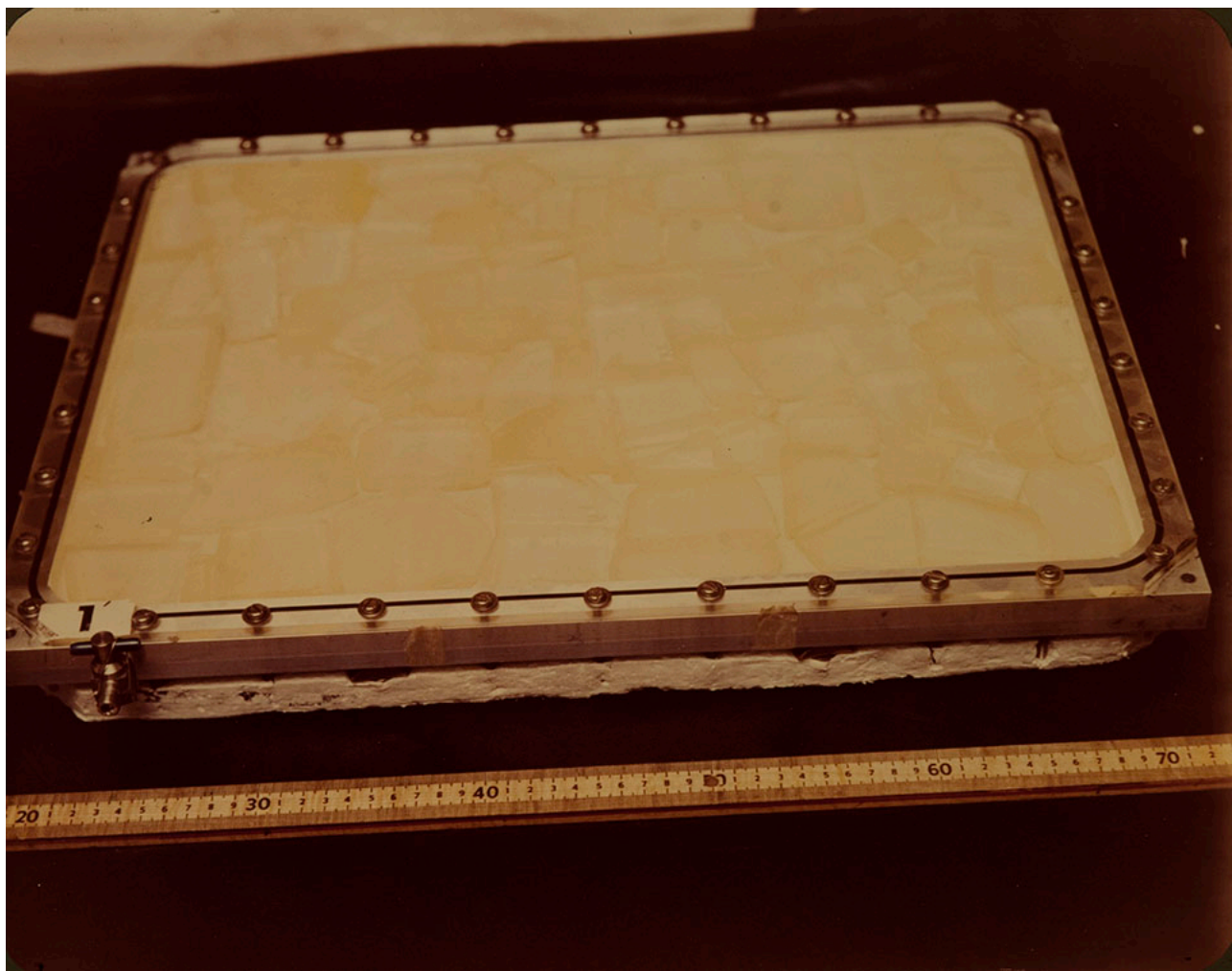
# **GRBs First Announced – 1973**

- **Ap. J. Preprint of Discovery**
- **AAS Meeting – Columbus, Ohio**

**Huntsville “Group” (2) - Decision to build a large-area  
detector balloon experiment**

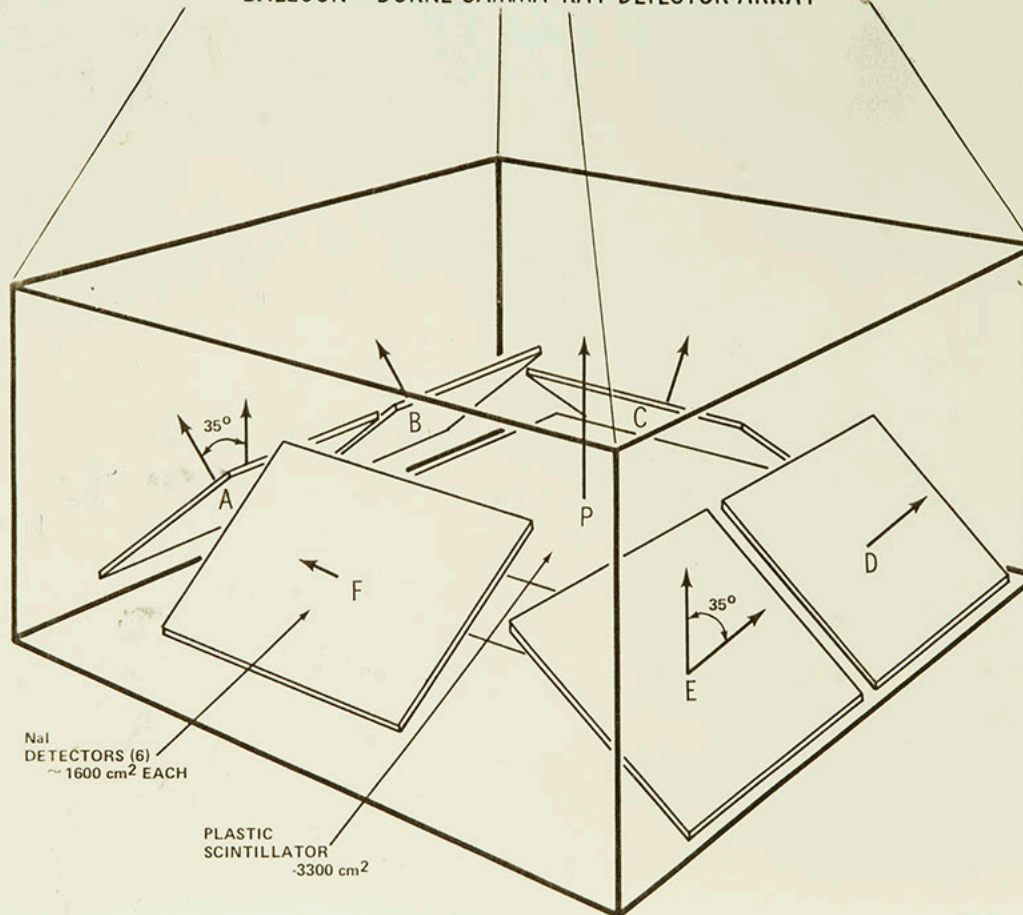
# **Detector Development & Early Balloon Flights**

**1974-1977**





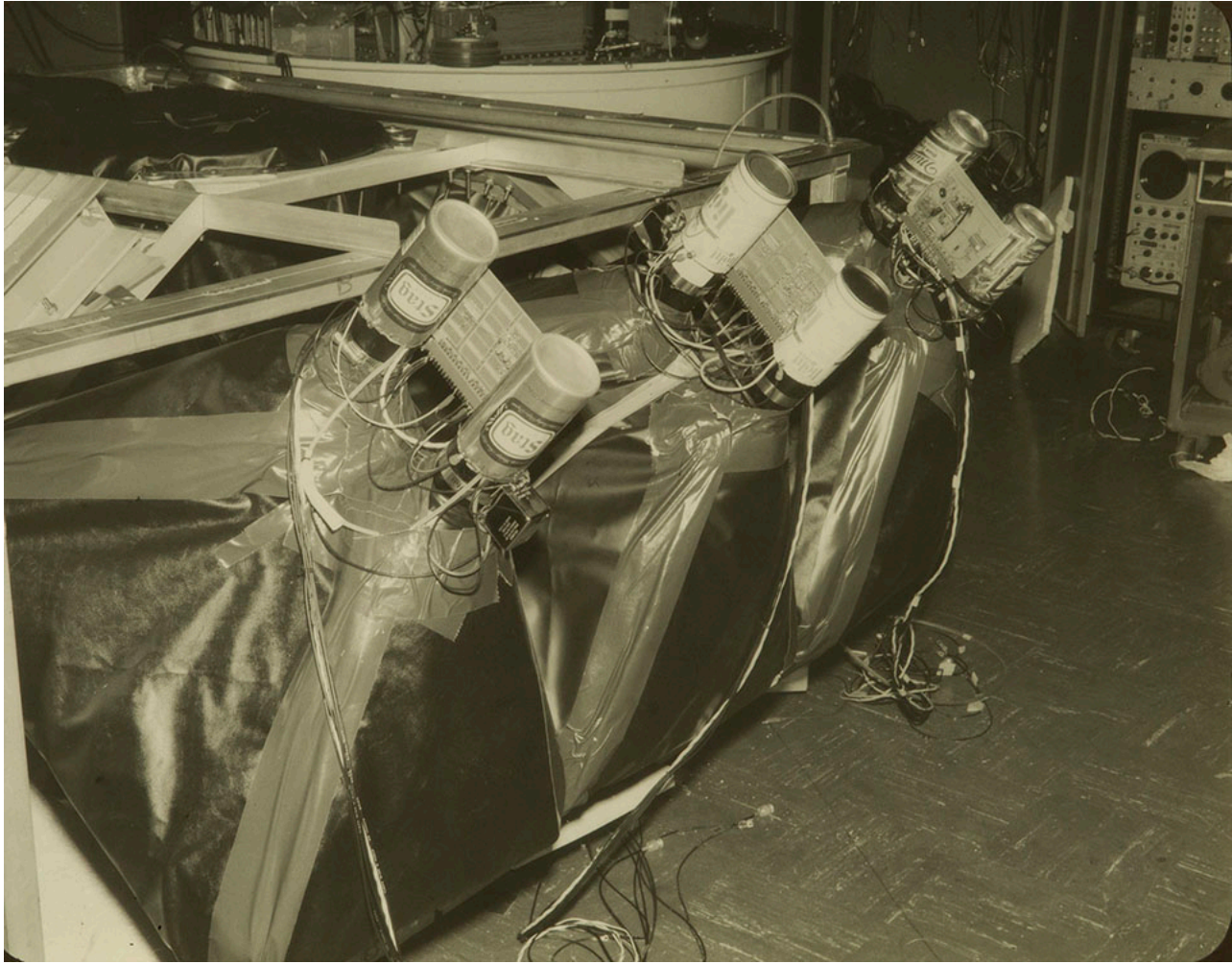
# BALLOON - BORNE GAMMA-RAY DETECTOR ARRAY



Nal  
DETECTORS (6)  
~ 1600 cm<sup>2</sup> EACH

PLASTIC  
SCINTILLATOR  
~ 3300 cm<sup>2</sup>





# 1977-1978

## GRO Announcement of Opportunity (AO)

- Proposed Transient Event Monitor (“TEM”)
- Twelve Detectors (Dodecahedron)
- Six on Top & Six on Bottom of S/C
- *Partial Selection* – Six Detectors
  - only on bottom of spacecraft



**1979-1980**

**“TEM” - Name changed to BATSE**

**Negotiating after Approval:**

- Requested 8 Detectors (for full-sky coverage)  
(Octahedron - Four on top & bottom)**

# GAMMA-RAY OBSERVATORY

## CONCEPTUAL DESIGN

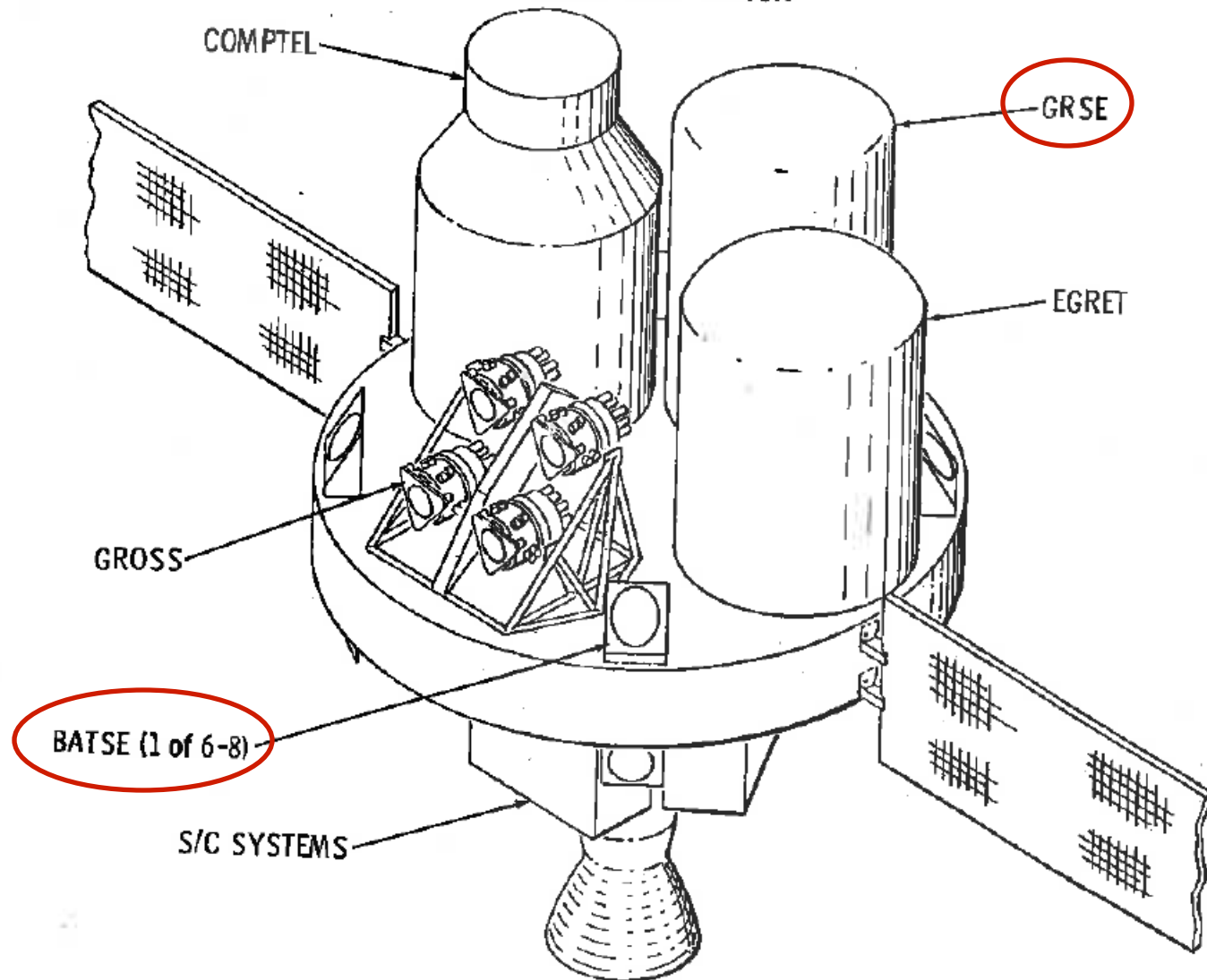
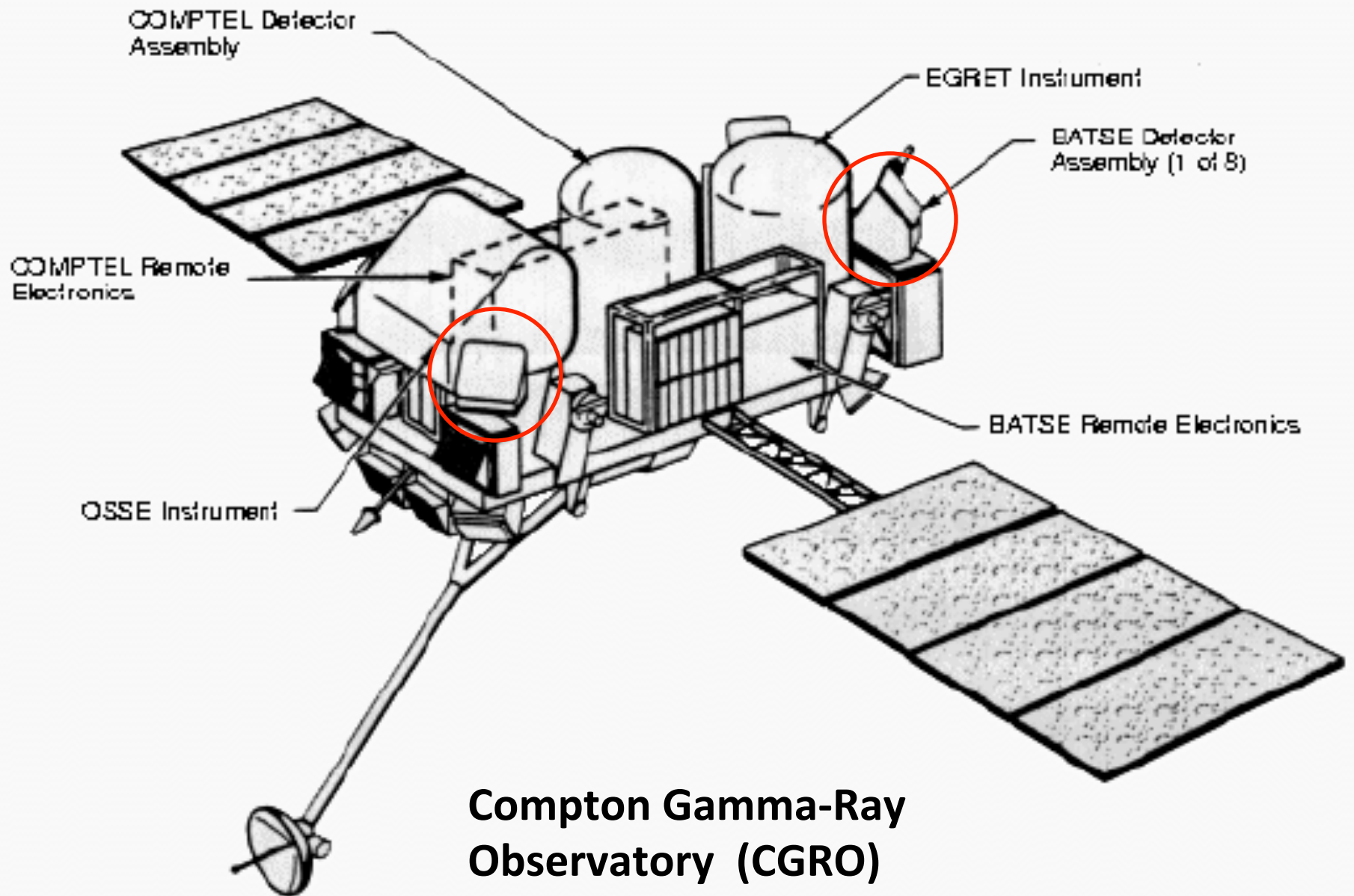


Fig. 3. Conceptual drawing of the planned Gamma-Ray Observatory (GRO) showing the four major pointed experiments and a possible configuration for the BATSE array.

# **1981-1982**

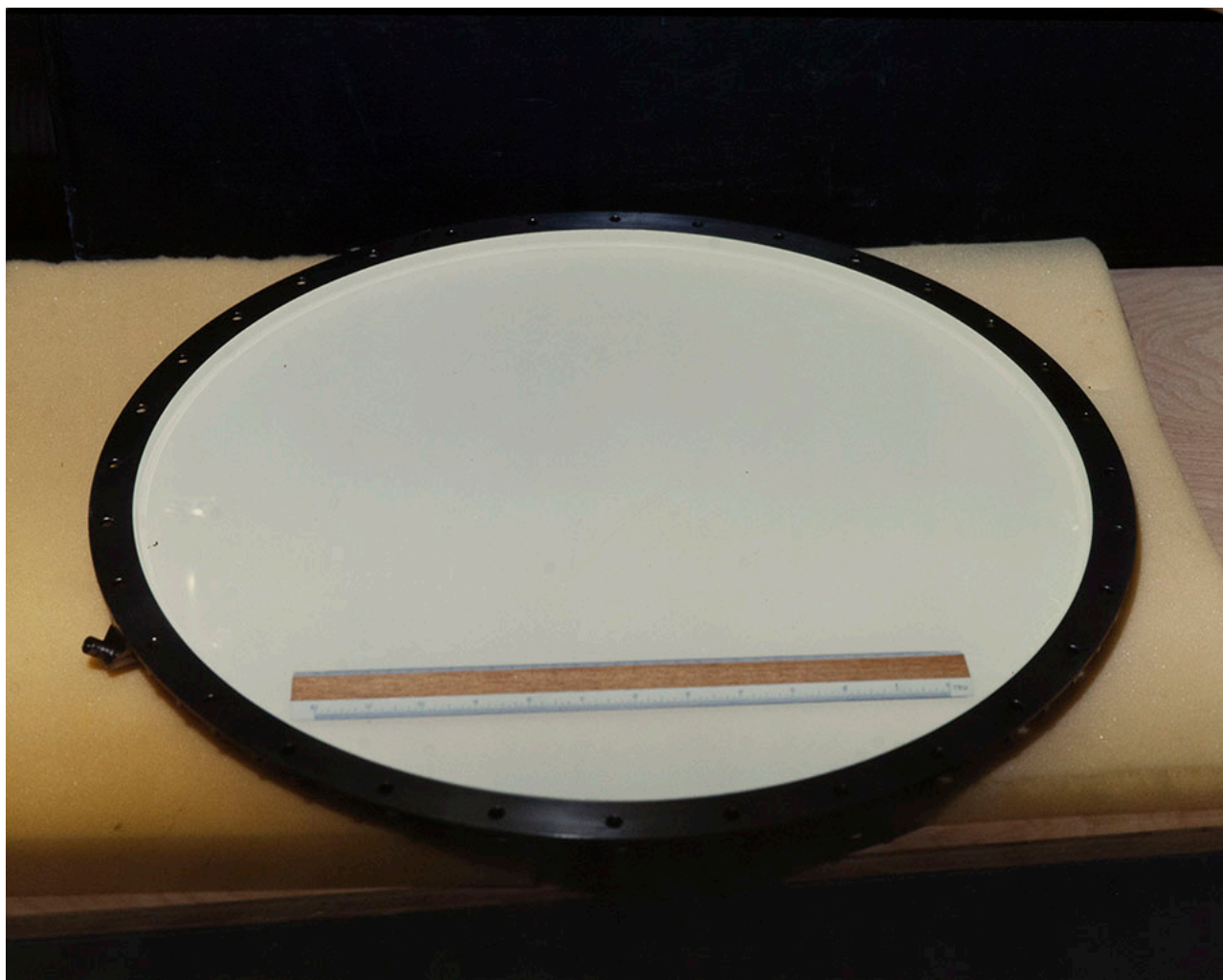
- GRSE Experiment Removed from GRO**
- Spectroscopy Detectors added to BATSE,  
one in each detector module**  
(a rare event; experiments are usually de-scoped by NASA)



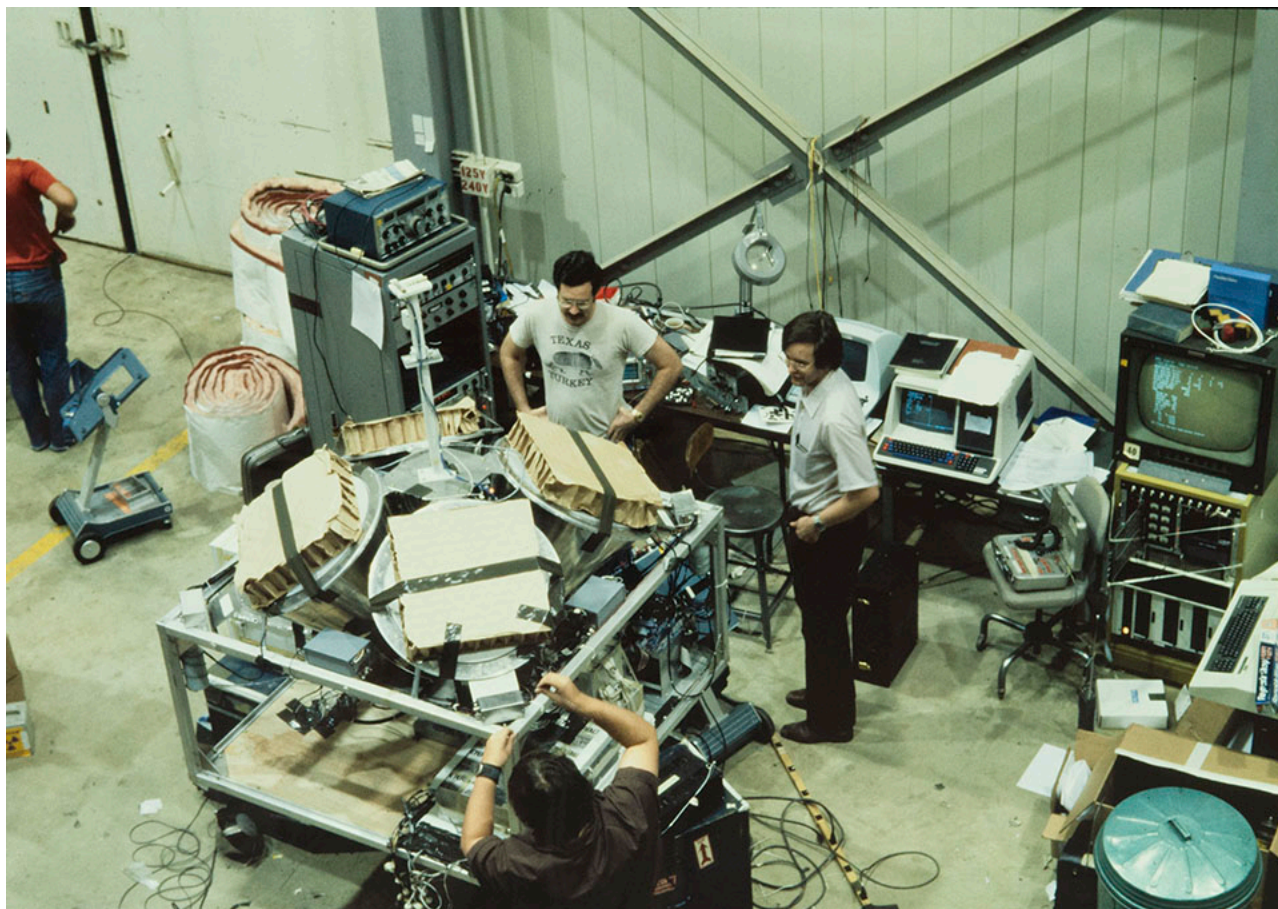
**Compton Gamma-Ray  
Observatory (CGRO)**

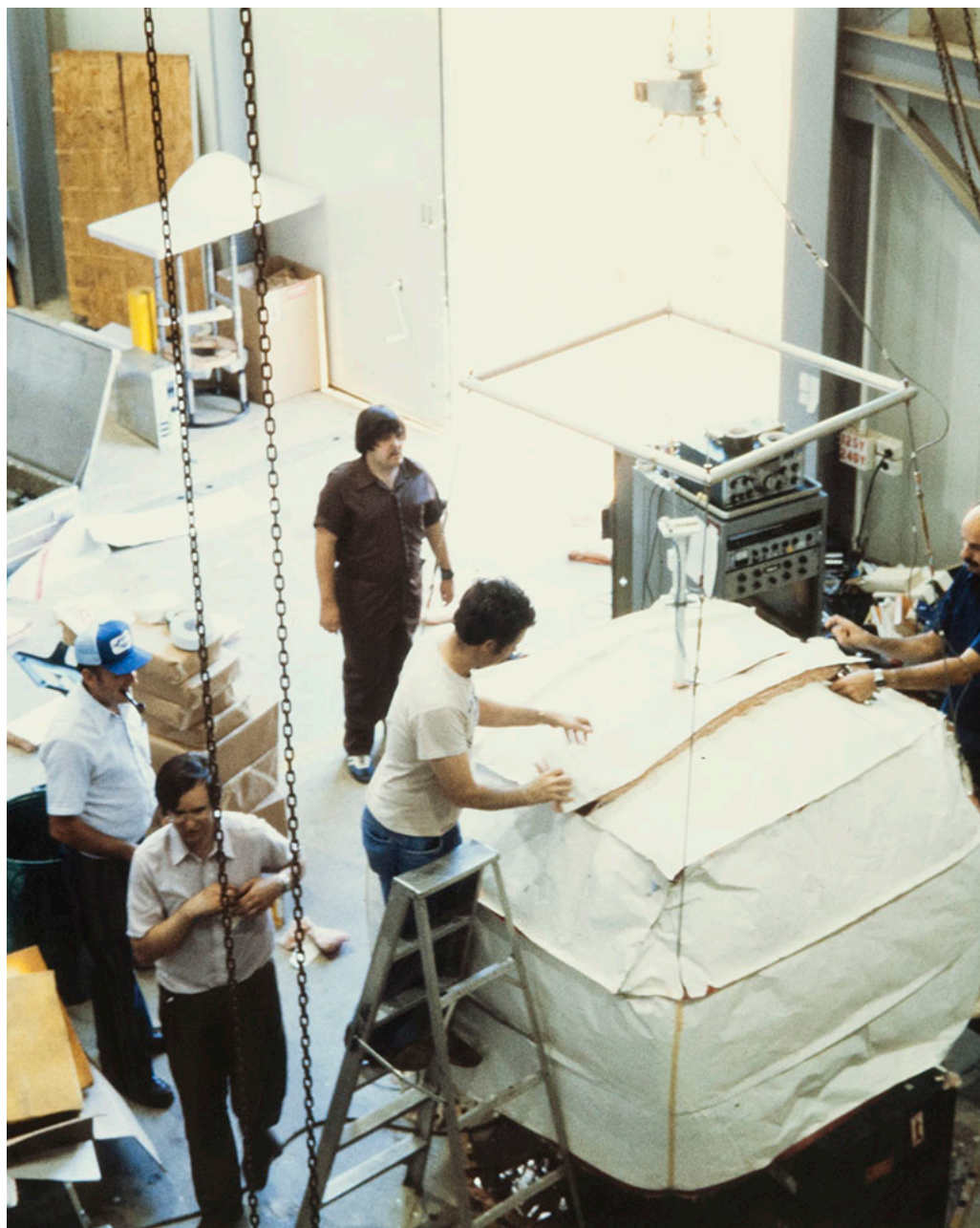
# **BATSE Detector Development & Balloon Flights**

**1981 - 1987**









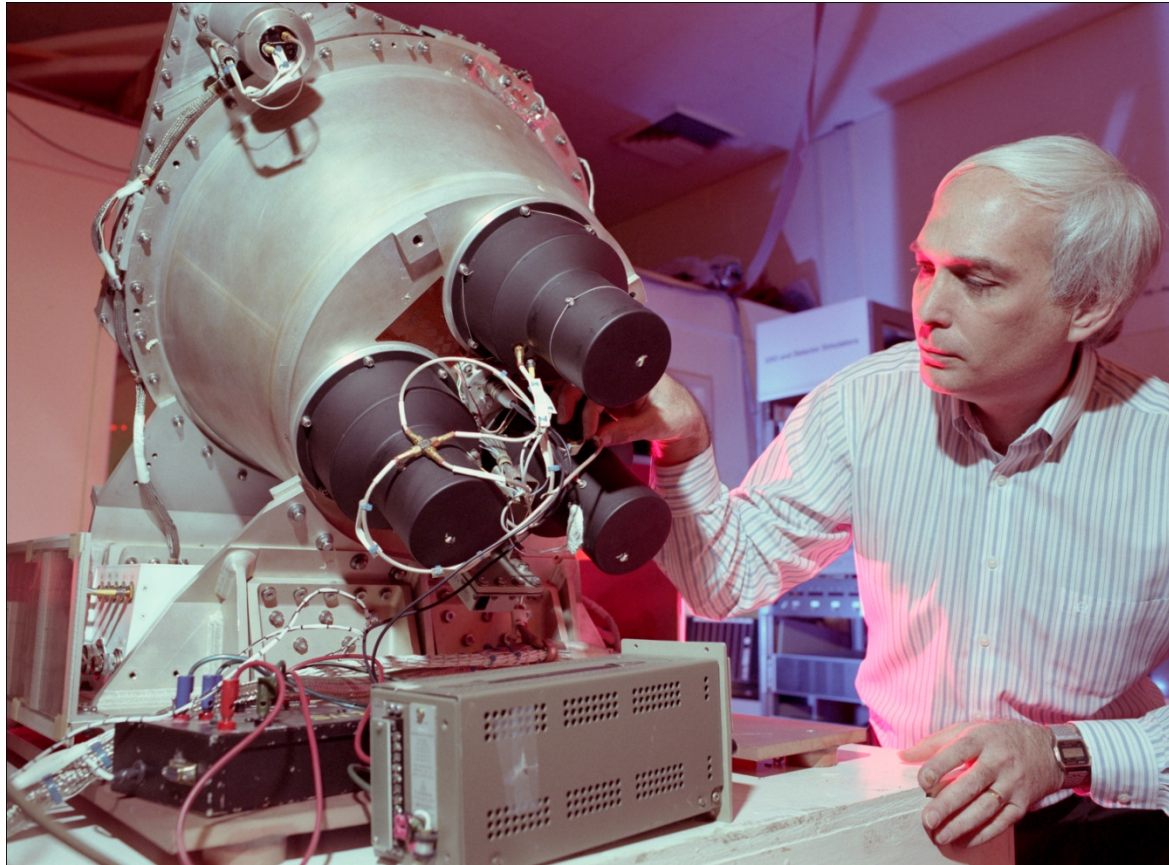




**Building BATSE & GRO**

**1982 - 1989**

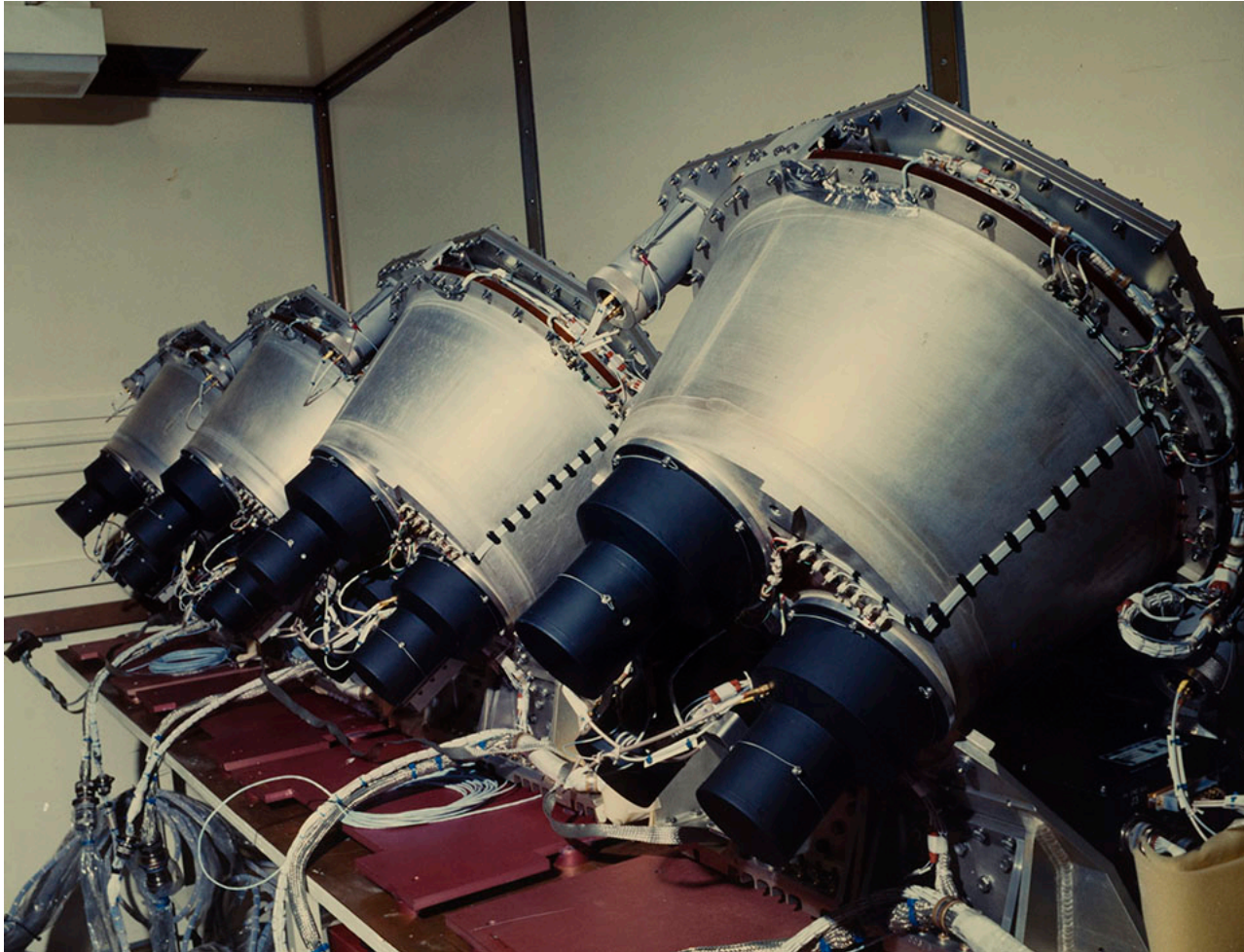
# BATSE Detector Module





# **BATSE Detector Modules:**

- Design, Fabrication, Testing & Calibration  
at NASA-MSFC**

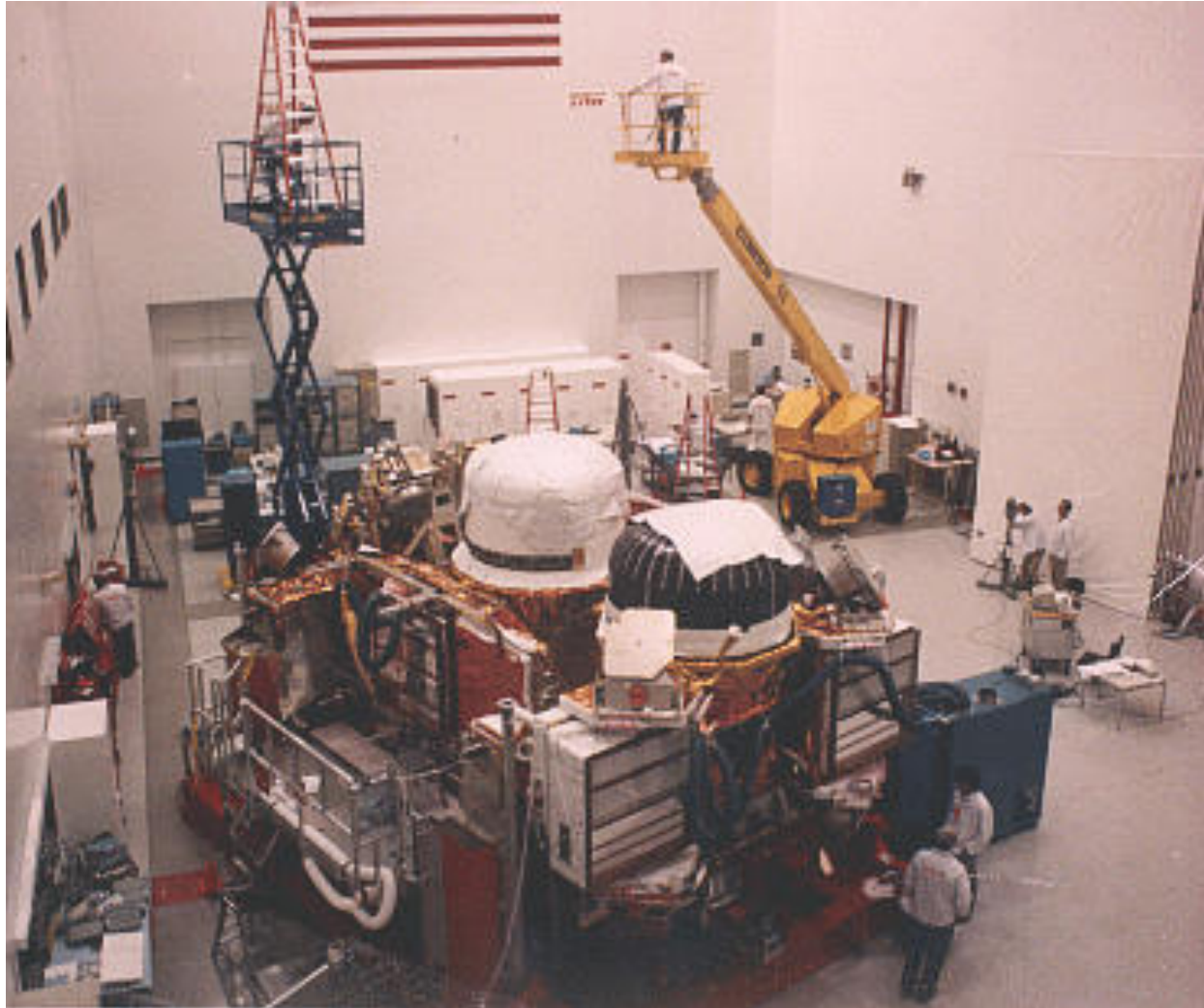




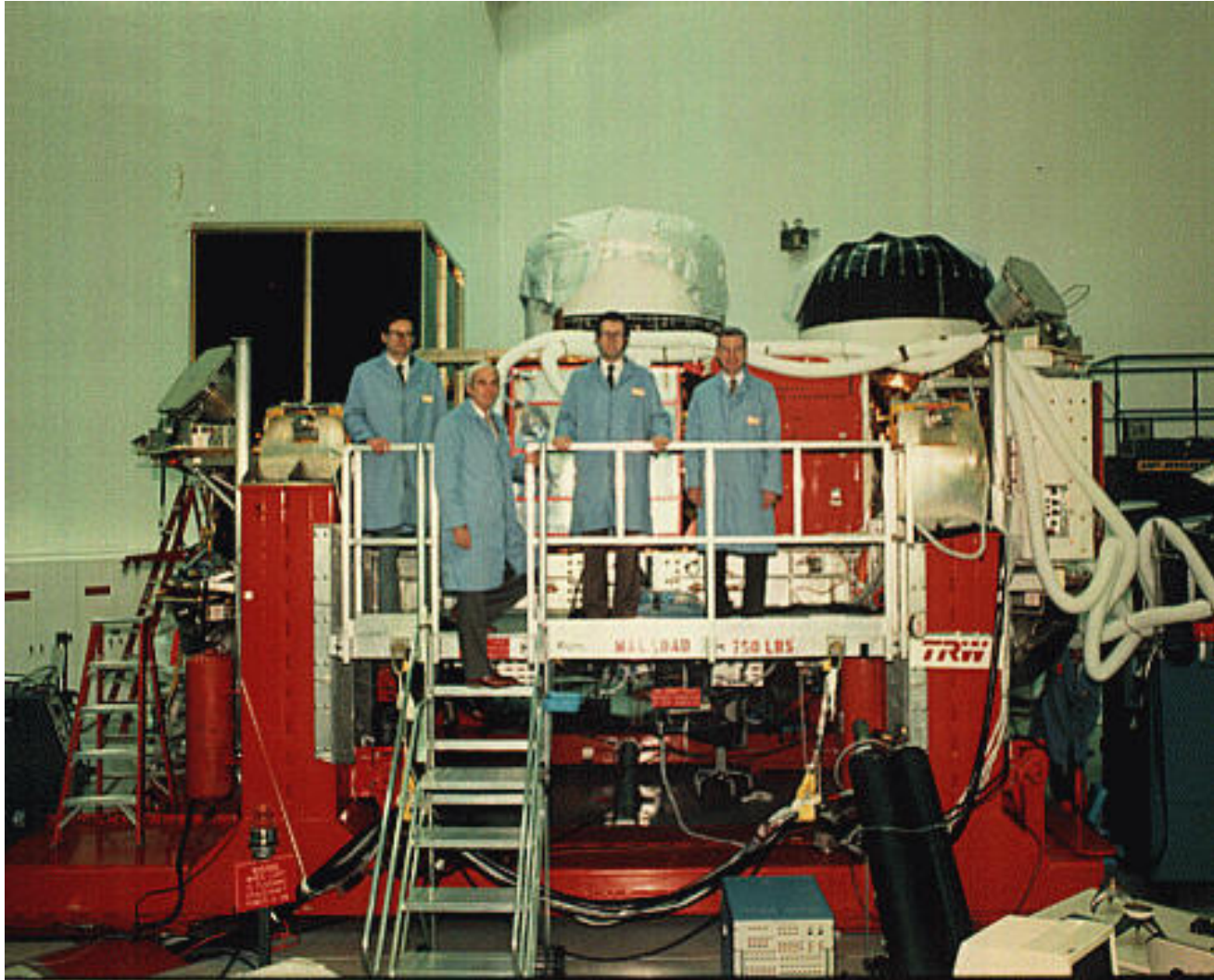
# Integration of Experiments on GRO Spacecraft in California



# Radioactive Source Survey of BATSE on the GRO Spacecraft



## GRO P.I.s Near the GRO Instruments



# **Launching & Deploying GRO**

**April 4-5, 1991**



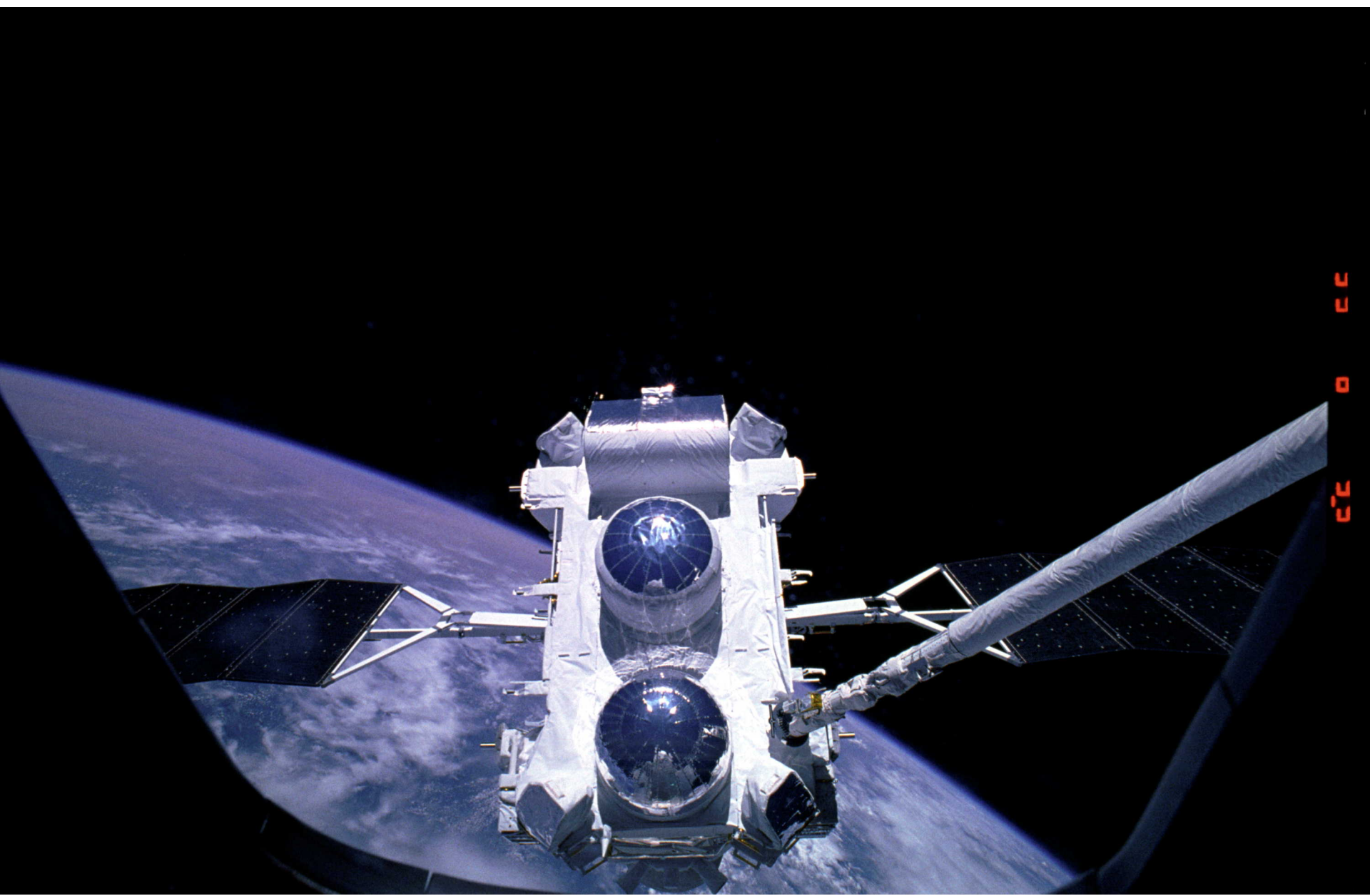
**Shuttle launch –  
Gamma-Ray  
Observatory (GRO)**

**(Operational:  
re-named CGRO)**

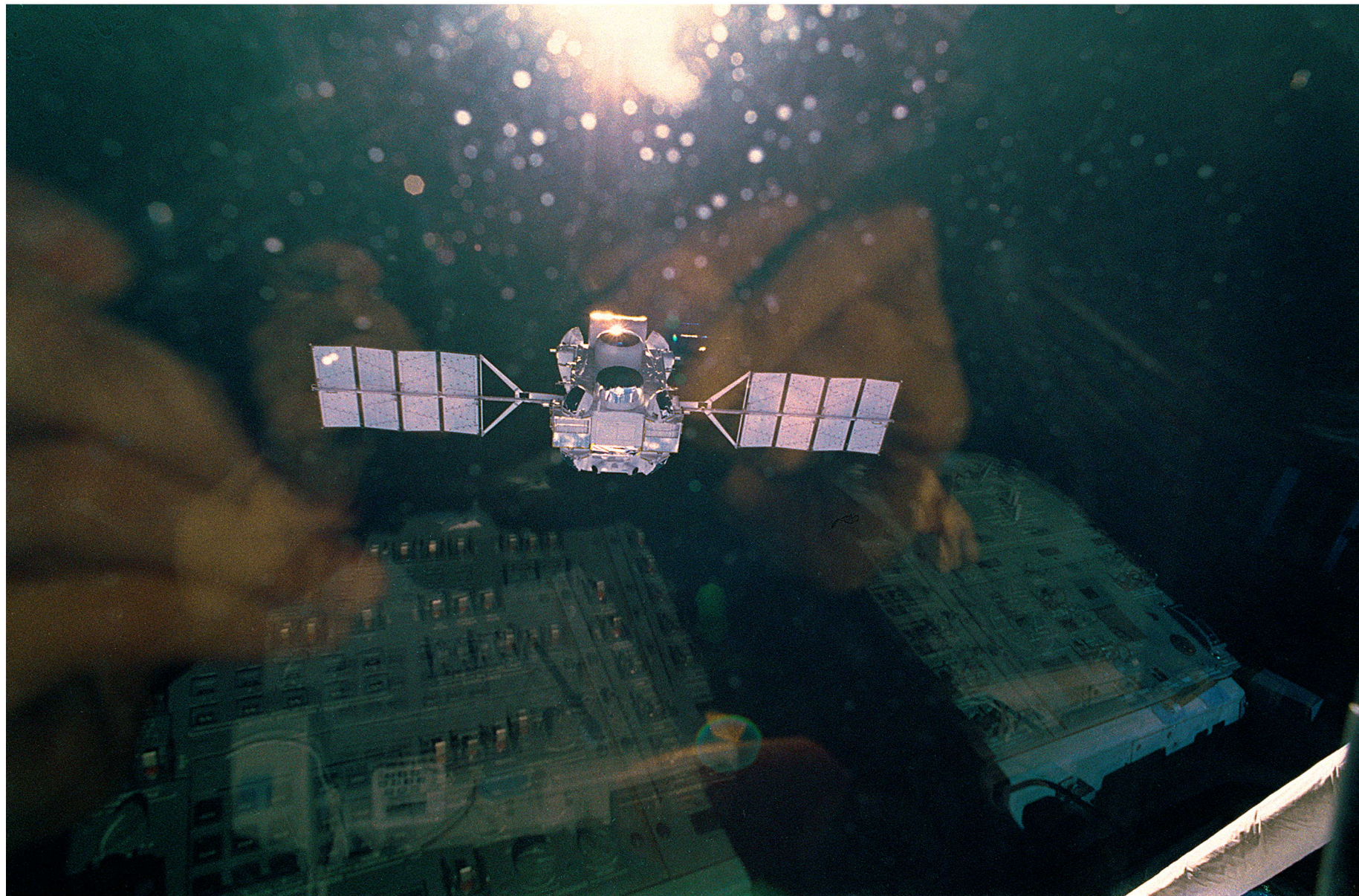














# nature

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## MAPPING COSMIC GAMMA-RAY BURSTS

An inventory for the origins of life  
More pulsar planets  
Atomic structure of bacteriophage  $\phi$ X174

# BATSE

**First large experiment designed for  
GRB studies:**

- **Full-sky**
- **Large area**
- **Good sensitivity for  $E_{\text{peak}}$  of most GRBs**
- **Moderate spectral capabilities** – (good enough for most time/spectral correlations and time resolved spectral studies)

# BATSE Legacy

- Largest sample (2704 GRBs); full-sky, 9+ years in operation
- Well- characterized, full-sky instrument
- Likely will not be exceeded for several decades
- Led to GCN Network (S. Barthelmy)

**>1000 Papers Based on BATSE Observations**

**>50 Ph.D. theses**

# **BATSE - Major GRB Results**

- **Global properties of GRB Distributions:**
  - Intensity Distribution & Sky Distribution
  - Not consistent with any Galactic Distribution, nearby extragal objects, incl. large clusters
  - Strong Indications that GRBs were at cosmological distances (Although BeppoSAX nailed it)
- **Comprehensive Temporal/Spectral Studies**
- **Two Populations of GRBs**
- **Rapid GRB Response:** beginning of Bacobine/GCN - led to breakthrough wide-field observations (e.g. GRB990123)



**The End**

# Back-up Notes:

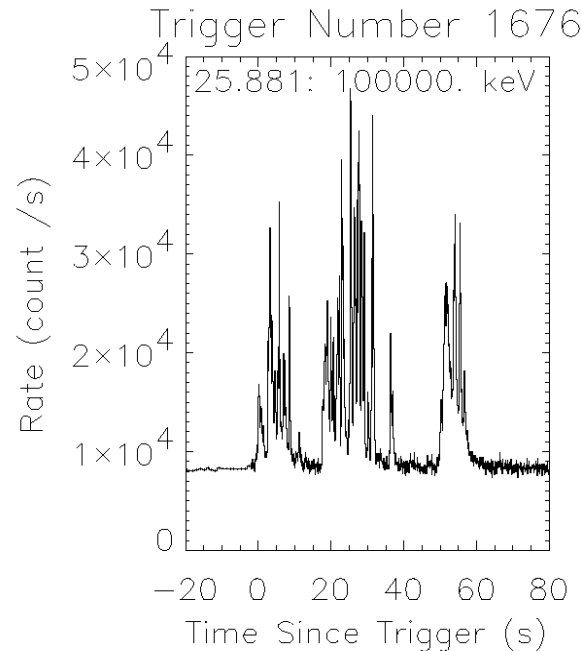
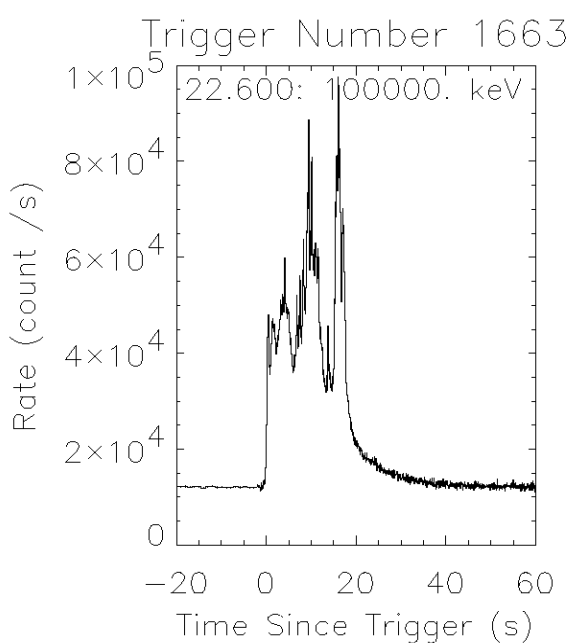
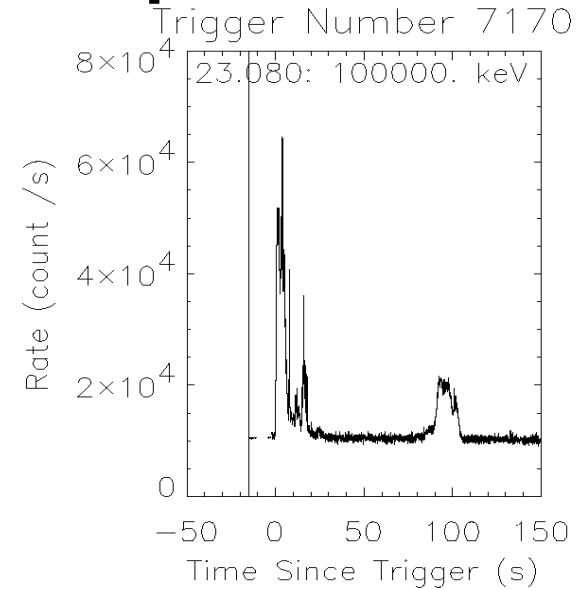
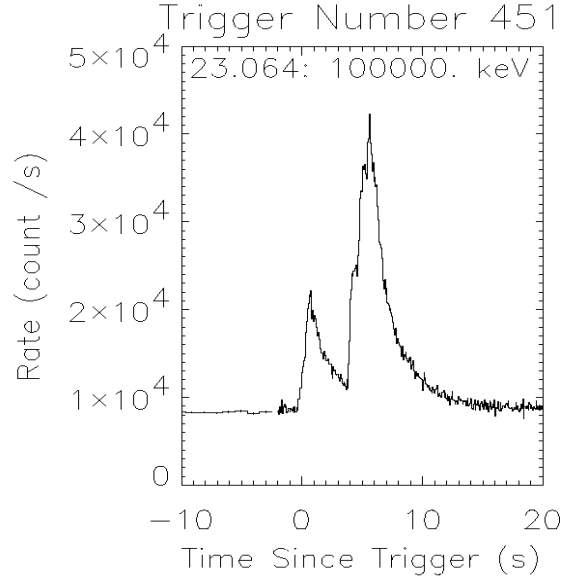
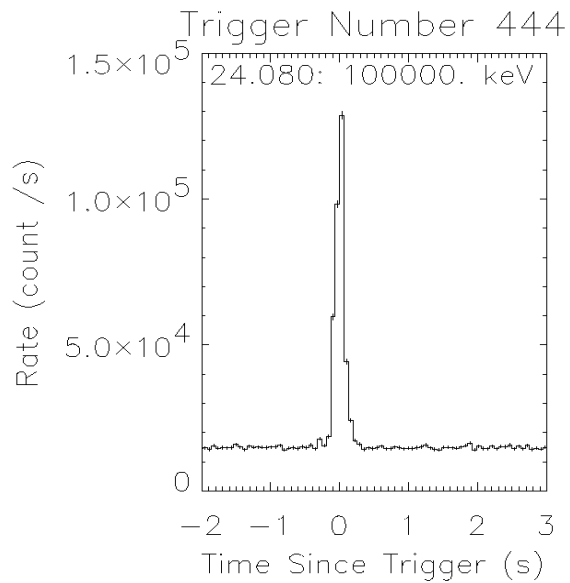
- First thoughts on GRB Observations:
  - Good field (nothing is known & high S/N )
- Need Large area to get many; remembered  $-3/2$  power law: will give ~several GRBs/day
- Low-cost balloon flights others had same idea
- - propose for GRO ( >>after Ltr. of Intent was due!

- Low cost; propose as signal to others.
- Propose dodecahedron (12) - 6 on top & 6 on bottom)
- Bottom was accepted; not the top, with 2 spares
- Fly the spares; need full sky w. same detectors; eight det. only (later a spare was approved

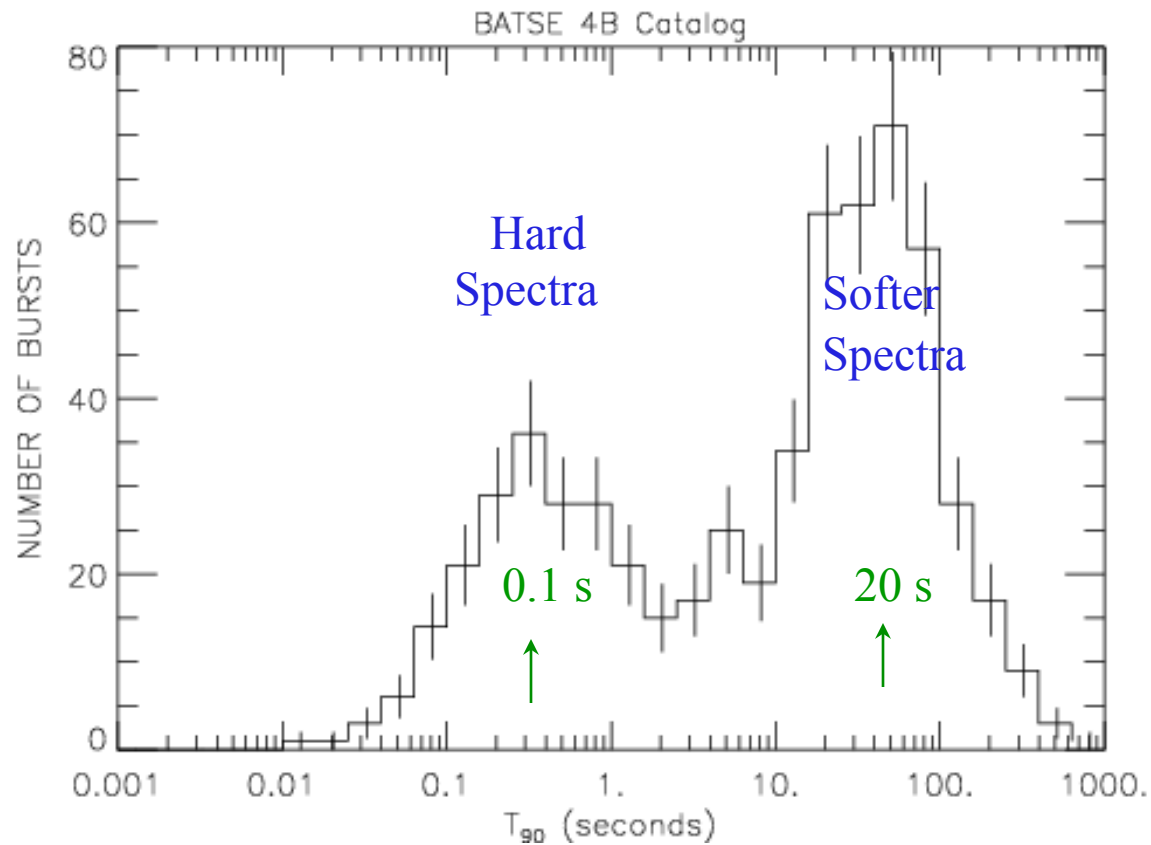
# **A Few Major BATSE GRB Results**

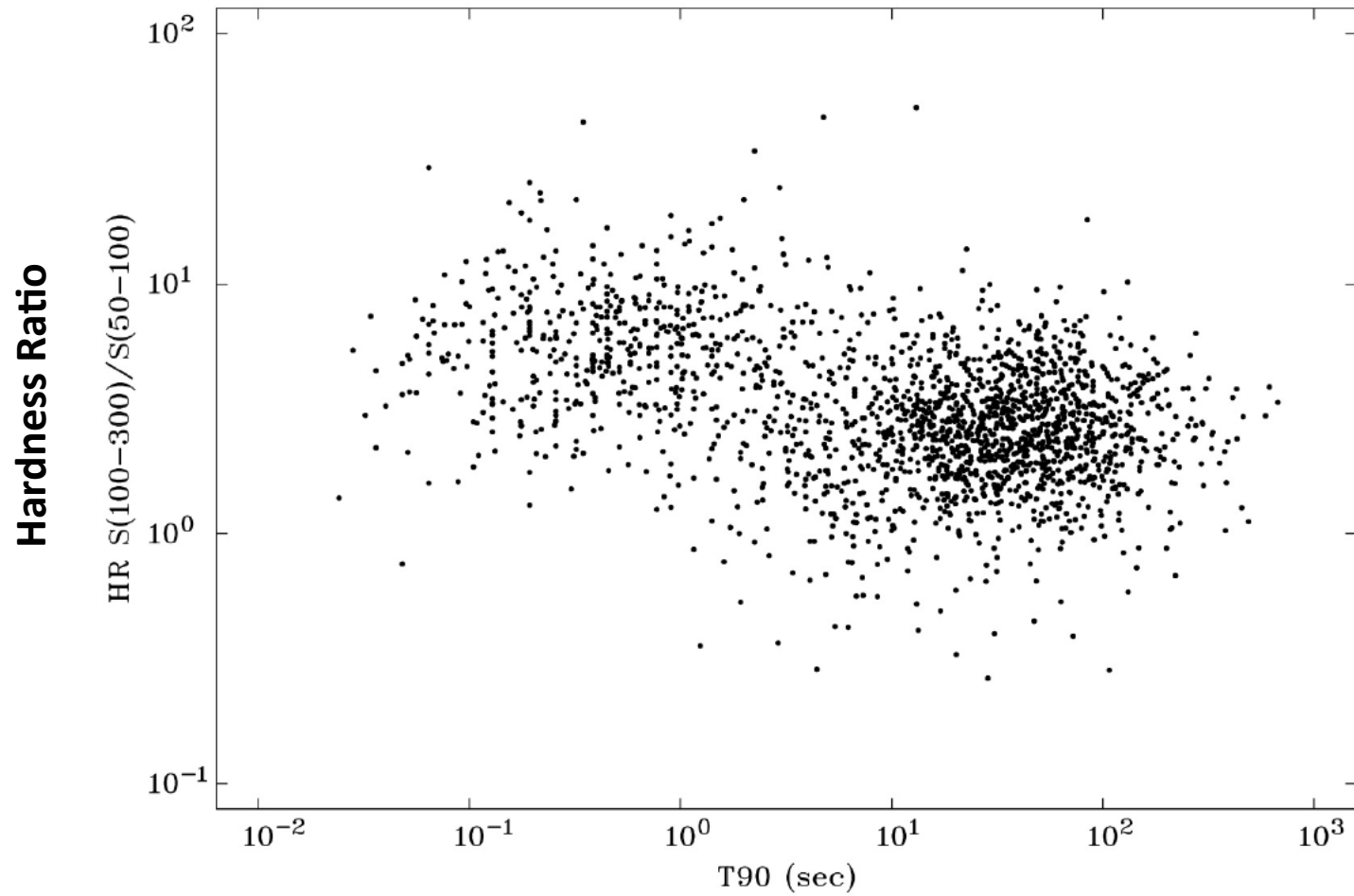
**1991-2000**

# Diversity of GRB Profiles & Coupled Spectral / Temporal Properties



## Two Distinct subclasses of $\gamma$ -ray bursts: short/hard & long/soft

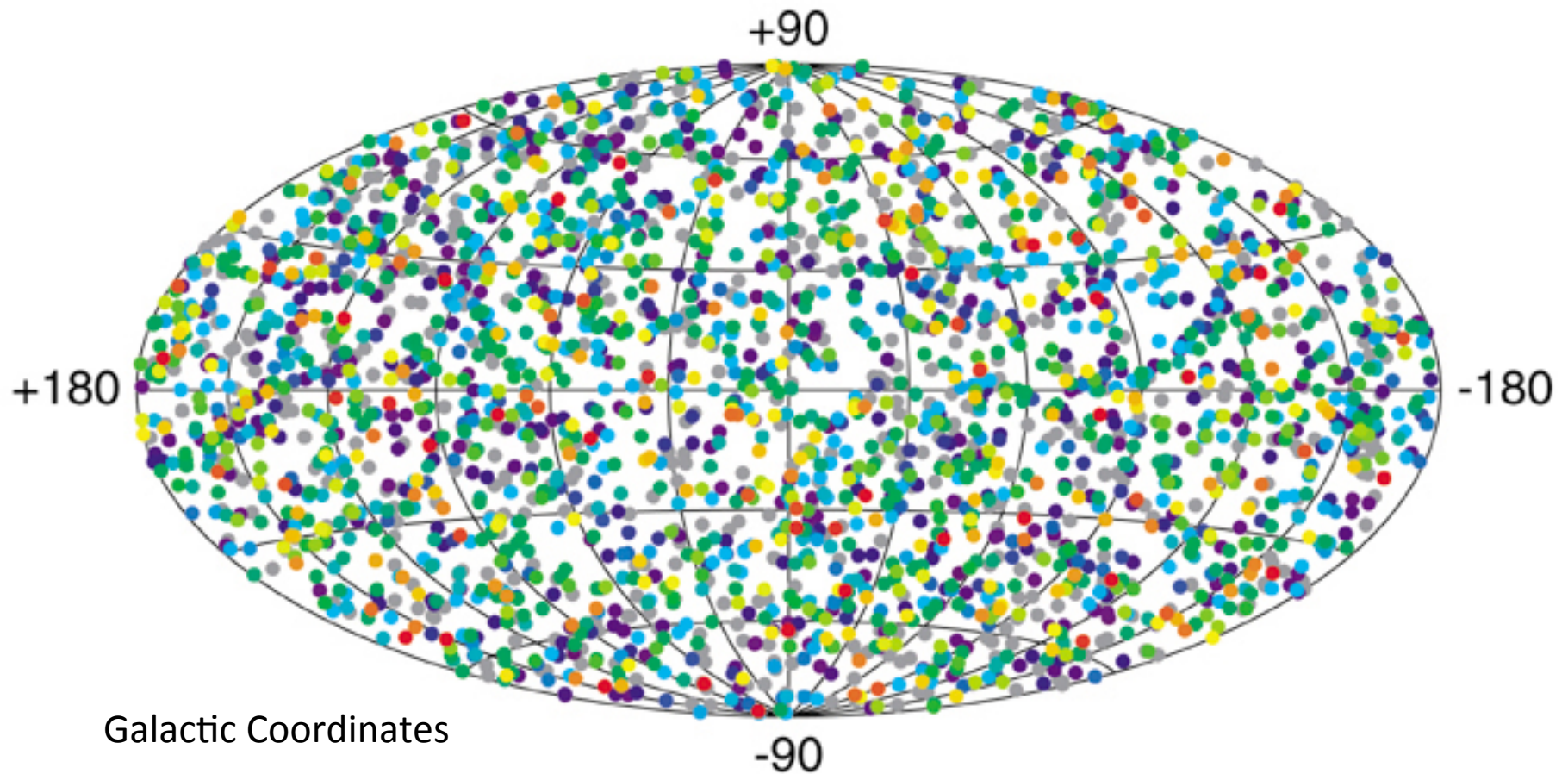




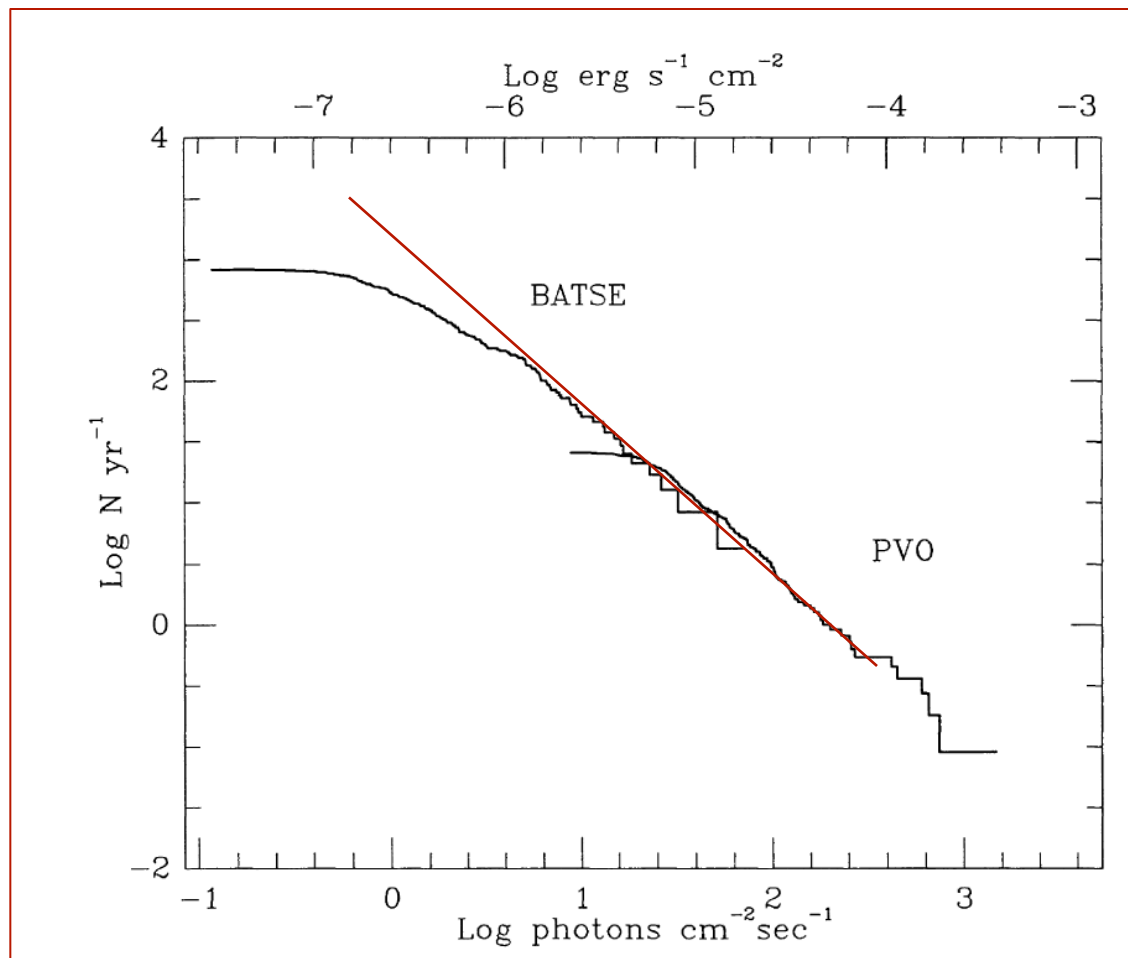
**Duration of Gamma-ray Bursts (sec)**

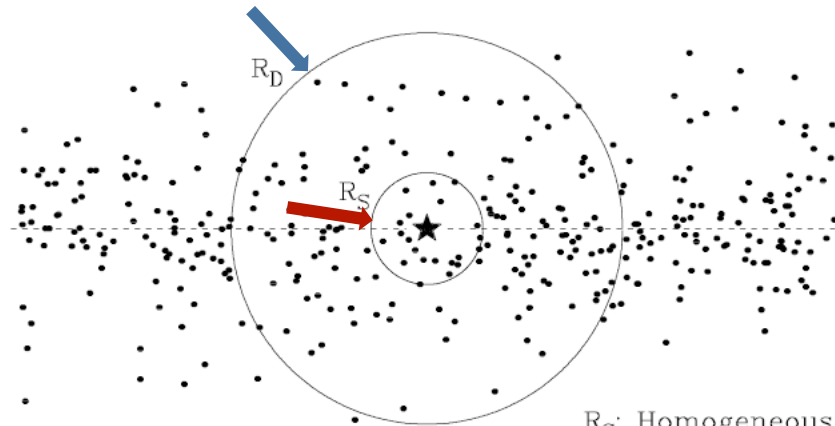


# 2704 BATSE Gamma-Ray Bursts

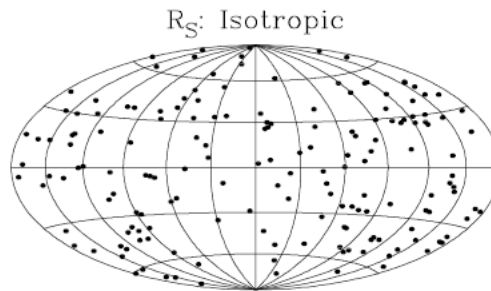


Apr. 1991 – May 2000

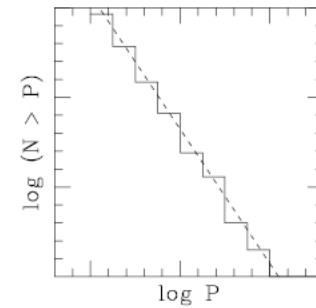




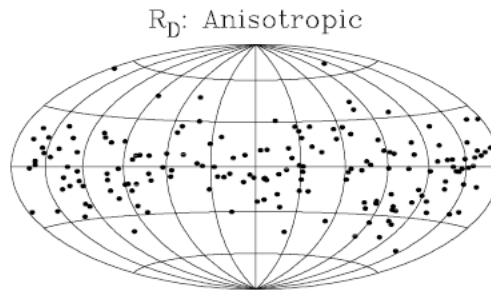
$R_S$  - Galactic, Nearby



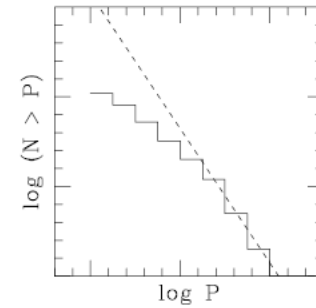
$R_S$ : Homogeneous

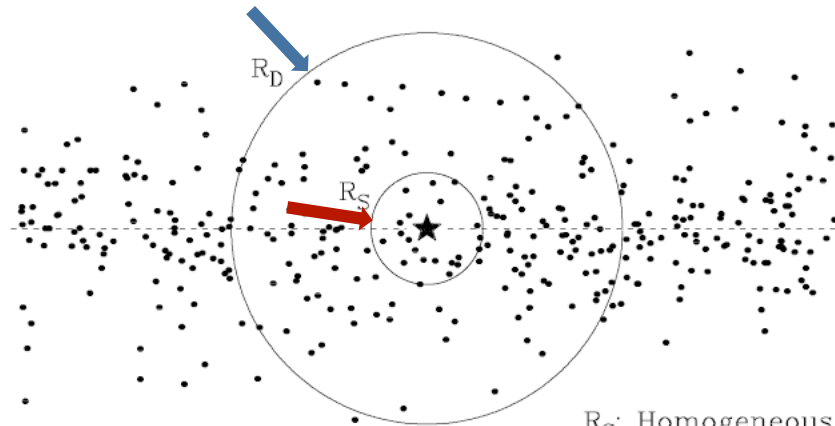


$R_D$  - Galactic, Far away



$R_D$ : Inhomogeneous

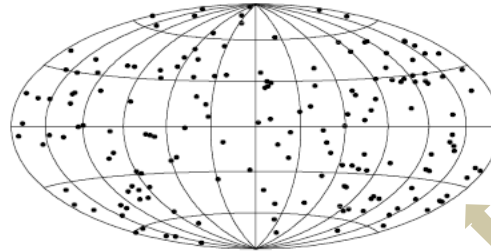




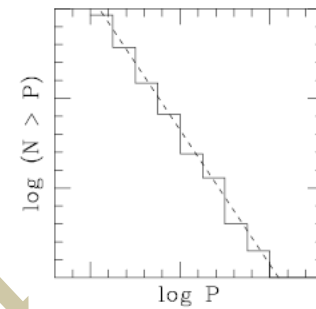
$R_S$  - Galactic, Nearby



$R_S$ : Isotropic



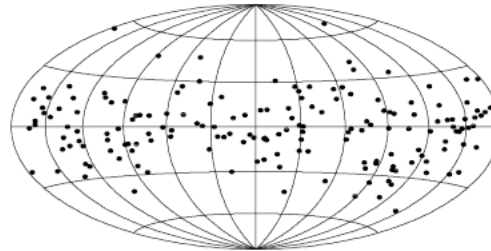
$R_S$ : Homogeneous



$R_D$  - Galactic, Far away



$R_D$ : Anisotropic



$R_D$ : Inhomogeneous

