Commutative Warm-up

NC Spacetime

Finding the Model

Properties of the Model

Conclusions and Outlook

sine-Gordon theory in noncommutative spacetime

S. Kürkçüoğlu

Institut für Theoretische Physik Leibniz Universität Hannover

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Outline				

- Commutative Warm-up: sine-Gordon model and a summary of well-known results.
- What is meant by "noncommutative" in this talk: Moyal algebra A_θ(ℝ^d) and ⋆-product.
- Finding the Model: Dimensional reduction from self-dual Yang-Mills(SDYM) theory.
- Properties of the Model: Classical and Quantum.
- Onclusions and Outlook.

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sine-Gordon Model				

Consider the following theory for a real scalar field in 1 + 1 dimensions.

$$S = \int dt dy \frac{1}{2} \partial_{\mu} \phi \partial^{\mu} \phi + 4 \alpha^2 (\cos \phi - 1).$$

- We use the metric η_{µν} = diag(1, -1), and α has the dimensions of mass.
- The equation of motion for ϕ is

$$\partial_{\mu}\partial^{\mu}\phi = -4\alpha^2 \sin\phi$$
.

 It has kink and anti-kink solutions, which are static and given by

$$\phi(\mathbf{y}) = \pm 4 \arctan e^{2\alpha \mathbf{y}}$$
.

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sine-Gordon Model				

It is energy density is localized. It is given by

$$\epsilon = \frac{1}{2}(\partial_y \phi)^2 + 4\alpha^2(1 - \cos \phi) = \frac{16\alpha^2}{\cosh^2 2\alpha y}$$

The kink and its energy density have the profiles



- Its classical mass is $M_{kink} = \int dy \epsilon = 16\alpha$.
- Kink has topological charge Q = 1. It is disconnected from the vacuum sector with Q = 0.

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sine-Gordon Model				

It is super-renormalizable. It is in fact integrable at the quantum level: Its S-matrix completely factorizes into two-particle S-matrices and obey Yang-Baxter equation. No particle production occurs!!!

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- It is super-renormalizable. It is in fact integrable at the quantum level: Its S-matrix completely factorizes into two-particle S-matrices and obey Yang-Baxter equation. No particle production occurs!!!
- It has an infinite set of conserved currents.

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sine-Gordon Model				

- It is super-renormalizable. It is in fact integrable at the quantum level: Its S-matrix completely factorizes into two-particle S-matrices and obey Yang-Baxter equation. No particle production occurs!!!
- It has an infinite set of conserved currents.
- It is equivalent to a fermionic theory, namely the massive Thirring model.

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sine-Gordon Model				

- It is super-renormalizable. It is in fact integrable at the quantum level: Its S-matrix completely factorizes into two-particle S-matrices and obey Yang-Baxter equation. No particle production occurs!!!
- It has an infinite set of conserved currents.
- It is equivalent to a fermionic theory, namely the massive Thirring model.
 - To explore indications of the model at quantum level, a simple analysis is to compute the corrections to *M_{kink}* by semi-classical means.

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sine-Gordon Model

Super-Renormalizable: It is sufficient to normal order the interactions to cancel all the divergences.

$$:4\alpha^{2}(\cos\phi-1):=4(\alpha^{2}-\delta\alpha^{2})(\cos\phi-1)$$

We can observe this quickly from the Feynman graphs. We have $\cos \phi - 1 = -\frac{1}{2}\phi^2 + \frac{1}{4!}\phi^4 - \frac{1}{6!}\phi^6 + \cdots$.

 All divergent contributions come from the self-contractions of the vertices.



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sine-Gordon Model				

Quantum Corrections to the Kink Mass

• This is done by finding the normal modes of the fluctuations around the kink solution. If ω_n are the frequencies of these modes, this implies

$$E_{kink-sector} = 16\alpha + \hbar \sum_{n} (k_n + \frac{1}{2})\omega_n + \nu_n + O(\alpha^2)$$

- $k_n = 0$ for the quantum kink particle, $k_n \neq 0$ for the scattering states of mesons in the presence of the kink particle.
- To find M_{kink} at this approximation, one subtracts E_{vacuum} and regularizes the remaining divergences by renormalizing α^2 . This gives

$$M_{kink} = 16lpha - rac{2}{\pi}lpha + O(lpha^2)$$

Noncommutative Spacetime: Moyal Algebra and *-product				
Definitions				
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Let A be the algebra of functions over ℝ^d. We multiply
 f, g ∈ A w.r.t. the pointwise multiplication map
 µ : A ⊗ A → A :

$$\mu(f(x)\otimes g(x))\equiv (f\cdot g)(x).$$

 Flat noncommutative spacetime is the associative algebra *A*_θ(ℝ^d)(Moyal Algebra) obtained by replacing μ with μ_{*}:

$$\mu_\star(f(x)\otimes g(x))\equiv (f\star g)(x)\,.$$

• *-product is given by the formula

$$(f\star g)(x) = f(x)e^{rac{i}{2} heta^{\mu
u}\overleftarrow{\partial_{\mu}}\overrightarrow{\partial_{\nu}}}g(x)$$

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Definitions				

 The coordinate functions x_μ generate A_θ(R^d) and they fulfil the commutation relation

$$\mathbf{X}_{\mu} \star \mathbf{X}_{\nu} - \mathbf{X}_{\nu} \star \mathbf{X}_{\mu} =: [\mathbf{X}_{\mu}, \mathbf{X}_{\nu}]_{\star} = i\theta_{\mu\nu} \,.$$

- $\theta_{\mu\nu}$ is a real antisymmetric tensor of rank 2, with constant components.
- In A_θ(R¹⁺¹) we will sometimes use the light-cone coordinates:

$$u = \frac{1}{2}(t+y), \quad v = \frac{1}{2}(t-y), \quad \partial_u = (\partial_t + \partial_y), \quad \partial_v = (\partial_t - \partial_y).$$

They fulfil

$$[\mathbf{v}, \mathbf{u}]_{\star} = i\theta$$
.

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We would like to have a NC sine-Gordon theory

Properties

- Classically Integrable: There is a linear system of equations, whose compatibility conditions implies a noncommutative version of sine-Gordon field equations.
- To have the correct commutative limit.
- To possess kink, anti-kink solutions.
- Causal S-matrix at tree-level.

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Further Properties?

- Semi-Classical behavior: Spectrum of quadratic fluctuations around the vacuum and kink solutions.
- Quantum corrections to the mass of the kink. Regularization of divergences.
- SUSY extensions and their properties.

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SDYM Theory and Dimensional Reduction					

- A well-known result is that SU(2) self-dual Yang-Mills (SDYM) theory on $R^{(2,2)}$ can be reduced to sine-Gordon model on $R^{(1,1)}$. Note that $R^{(2,2)}$ has signature (-++-).
- So consider the self-dual U(2) SDYM on $\mathcal{A}_{\theta}(\mathbb{R}^{(2,2)})$. (We follow Lechtenfeld et. al. *Nucl.Phys.B***705**(2005))

$$F_{\mu\nu} = \frac{1}{2} \varepsilon_{\mu\nu\rho\sigma} F^{\rho\sigma}, \quad F_{\mu\nu} = \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu} + [A_{\mu}, A_{\nu}]_{\star}$$

- First we will reduce to a theory in $\mathcal{A}_{\theta}(\mathbb{R}^{(2,1)})$.
- We take $A_4 = \Lambda$ and demand translational invariance of A_{μ} along x_4 -direction: i.e. $\partial_4(A_a, A_4) = 0$. This gives:

$$\partial_a \Lambda + [A_a, \Lambda]_{\star} = \frac{1}{2} \varepsilon_{abc} F^{bc}.$$

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Gauge Fixing

We fix the gauge by

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$$\begin{array}{rcl} A_t - A_y &=& 0\,, & A_t + A_y = \Phi^{-1} \star (\partial_t + \partial_y) \Phi\,, \\ A_x + \Lambda &=& 0\,, & A_x + \Lambda = \Phi^{-1} \star \partial_x \Phi\,. \end{array}$$

• Here Φ is a U(2) valued field.

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Gauge Fixing

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$$\begin{aligned} & A_t - A_y &= 0, \quad A_t + A_y = \Phi^{-1} \star (\partial_t + \partial_y) \Phi \\ & A_x + \Lambda &= 0, \quad A_x + \Lambda = \Phi^{-1} \star \partial_x \Phi \,. \end{aligned}$$

• Here Φ is a U(2) valued field.

Introducing the light-cone coordinates

$$u = \frac{1}{2}(t+y), \quad v = \frac{1}{2}(t-y), \quad \partial_u = (\partial_t + \partial_y), \quad \partial_v = (\partial_t - \partial_y).$$

We find

$$\partial_{\chi}(\Phi^{-1}\star\partial_{\chi}\Phi) - \partial_{\nu}(\Phi^{-1}\star\partial_{\mu}\Phi) = 0.$$

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A Linear System				
Linear Svs	tom			

• Consider the system of equations

$$(\zeta \partial_x - \partial_u)\Psi = \Phi^{-1} \star \partial_u \Phi \star \Psi, \quad (\zeta \partial_v - \partial_x)\Psi = \Phi^{-1} \star \partial_x \Phi \star \Psi$$

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A Lincor System				

Linear System

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look

• $\Psi(x, u, v, \zeta)$ is valued in U(2) and $\zeta \in \mathbb{C}P^1$

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A Linear System

Linear System

Consider the system of equations

$$(\zeta \partial_x - \partial_u)\Psi = \Phi^{-1} \star \partial_u \Phi \star \Psi, \quad (\zeta \partial_v - \partial_x)\Psi = \Phi^{-1} \star \partial_x \Phi \star \Psi$$

- $\Psi(x, u, v, \zeta)$ is valued in U(2) and $\zeta \in \mathbb{C}P^1$
- There is the reality condition $\Psi(\cdot, \zeta) \star \Psi^{\dagger}(\cdot, \overline{\zeta}) = 1$.

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A Linear System

Linear System

Consider the system of equations

$$(\zeta \partial_x - \partial_u)\Psi = \Phi^{-1} \star \partial_u \Phi \star \Psi, \quad (\zeta \partial_v - \partial_x)\Psi = \Phi^{-1} \star \partial_x \Phi \star \Psi$$

- $\Psi(x, u, v, \zeta)$ is valued in U(2) and $\zeta \in \mathbb{C}P^1$
- There is the reality condition $\Psi(\cdot, \zeta) \star \Psi^{\dagger}(\cdot, \overline{\zeta}) = 1$.
- We further have $\Psi(\cdot, \zeta \to 0) = \Phi^{-1}$.

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A Linear System

Linear System

Consider the system of equations

$$(\zeta \partial_x - \partial_u)\Psi = \Phi^{-1} \star \partial_u \Phi \star \Psi, \quad (\zeta \partial_v - \partial_x)\Psi = \Phi^{-1} \star \partial_x \Phi \star \Psi$$

- $\Psi(x, u, v, \zeta)$ is valued in U(2) and $\zeta \in \mathbb{C}P^1$
- There is the reality condition $\Psi(\cdot, \zeta) \star \Psi^{\dagger}(\cdot, \overline{\zeta}) = \mathbf{1}$.
- We further have $\Psi(\cdot, \zeta \to 0) = \Phi^{-1}$.

Compatibility Condition

Compatibility condition for this linear system is

$$\partial_x(\Phi^{-1}\star\partial_x\Phi)-\partial_v(\Phi^{-1}\star\partial_u\Phi)=0.$$

- This is in fact the NC version of integrable Ward model.
- It is reminiscent to NC version of nonlinear sigma model and it does have multi-soliton solutions.

Conclusions and Outlook

Reduction to 1 + 1-Dimensions

In two steps: First we factorize *x*-dependence, then restrict the form of the U(2) matrices.

- We take $[t, x]_{\star} = 0$, $[x, y]_{\star} = 0$, $[t, y]_{\star} = i\theta$.
- Take the ansatz

 $\Phi(t,x,y) = V(x) g(t,y) V^{\dagger}(x), V(x) = e^{i\alpha x\sigma_1}, g(t,y) \in U(2).$

Linear system becomes

$$\begin{array}{lll} \partial_u \Psi - i\alpha \zeta[\sigma_1\,,\Psi] &=& -V^{-1}g^{-1} \star \partial_u gV \star \Psi\,,\\ \zeta \partial_v \Psi - i\alpha[\sigma_1\,,\Psi] &=& V^{-1}g^{-1} \star \partial_x gV \star \Psi\,. \end{array}$$

• We have some freedom to pick $g(t, y) \in U(2)$. We choose

$$egin{array}{ccc} g=\left(egin{array}{ccc} g_+ & 0\ 0 & g_- \end{array}
ight) & \in U(1)\otimes U(1)\subset U(2) \end{array}$$

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Reduction to 1 + 1-Dimensions

Compatibility condition implies the equations

NC sine-Gordon equations

$$\partial_{\nu}(g_{+}^{-1} \star \partial_{u}g_{+}) + \alpha^{2}(g_{-}^{-1} \star g_{+} - g_{+}^{-1} \star g_{-}) = 0$$

$$\partial_{\nu}(g_{-}^{-1} \star \partial_{u}g_{-}) + \alpha^{2}(g_{+}^{-1} \star g_{-} - g_{-}^{-1} \star g_{+}) = 0$$

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Reduction to 1 + 1-Dimensions

Compatibility condition implies the equations

NC sine-Gordon equations

$$\partial_{\nu}(g_{+}^{-1} \star \partial_{u}g_{+}) + \alpha^{2}(g_{-}^{-1} \star g_{+} - g_{+}^{-1} \star g_{-}) = 0$$

$$\partial_{\nu}(g_{-}^{-1} \star \partial_{u}g_{-}) + \alpha^{2}(g_{+}^{-1} \star g_{-} - g_{-}^{-1} \star g_{+}) = 0$$

• It is possible to parameterize g_{\pm} by

$$g_+=e_\star^{-i\phi_+}\,,\quad g_-=e_\star^{i\phi_-}$$

• Taking $\theta \to 0$ and using $\varphi := \phi_+ + \phi_-$ and $\rho := \phi_+ - \phi_-$, leads to

$$\partial_{u}\partial_{v}\varphi = -4\alpha^{2}\sin\varphi, \quad \partial_{u}\partial_{v}\rho = 0.$$

 Thus we propose the equations above as the field equations for the NC sine-Gordon model.

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The Action				
If $lpha$ was 0, we would have had				

 $\partial_{\nu}(g_+^{-1}\star\partial_u g_+)=0\,,\quad \partial_{\nu}(g_-^{-1}\star\partial_u g_-)=0\,.$

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	The Action						
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$\partial_{\nu}(g_+^{-1}\star\partial_u g_+)=0\,,\quad \partial_{\nu}(g_-^{-1}\star\partial_u g_-)=0\,.$

• These type of equations are typical of WZW models.

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The Action				
If $lpha$ was 0, we would have had				

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$$\partial_{\boldsymbol{\nu}}(g_+^{-1}\star\partial_{\boldsymbol{u}}g_+)=0\,,\quad\partial_{\boldsymbol{\nu}}(g_-^{-1}\star\partial_{\boldsymbol{u}}g_-)=0\,.$$

- These type of equations are typical of WZW models.
- The action should be consisting of WZW action for g₊ and g₋, plus an interaction term:

Action

 $S[g_+,g_-] = S_{WZW}[g_+] + S_{WZW}[g_-] + \alpha^2 \int dt dy \, (g_+^{\dagger} \star g_- + g_-^{\dagger} \star g_+ - 2) \, .$

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The Action				
If $lpha$ was 0, we would have had				

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$$\partial_{\boldsymbol{\nu}}(g_+^{-1}\star\partial_{\boldsymbol{u}}g_+)=0\,,\quad\partial_{\boldsymbol{\nu}}(g_-^{-1}\star\partial_{\boldsymbol{u}}g_-)=0\,.$$

- These type of equations are typical of WZW models.
- The action should be consisting of WZW action for g₊ and g₋, plus an interaction term:

Action

$$\begin{split} S[g_+,g_-] &= S_{WZW}[g_+] + S_{WZW}[g_-] + \alpha^2 \int dt dy \, (g_+^{\dagger} \star g_- + g_-^{\dagger} \star g_+ - 2) \\ S_{WZW}[f] &= -\frac{1}{2} \int dt dy \, \partial_{\mu} f^{-1} \star \partial^{\mu} f \\ &- \frac{1}{3} \int dt dy \int_0^1 d\lambda \varepsilon^{\mu\nu\sigma} \, \hat{f}^{-1} \partial_{\mu} \hat{f} \star \hat{f}^{-1} \partial_{\nu} \hat{f} \star \hat{f}^{-1} \partial_{\sigma} \hat{f} \,. \end{split}$$

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The model has the standard static kink, anti-kink solutions.

Kink, Anti-Kink

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$$arphi_0=\pm 4\, {
m arctan}\, e^{2lpha y}\,,\quad
ho_0=0\,,\quad g_0=e^{-rac{i}{2}arphi_0}\,,$$

 Multi-soliton configurations can be constructed using the linear system via the "dressing" method. Commutative Warm-up

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The model has the standard static kink, anti-kink solutions.

Kink, Anti-Kink

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$$\varphi_0 = \pm 4 \arctan e^{2lpha y}, \quad \rho_0 = 0, \quad g_0 = e^{-\frac{i}{2}\varphi_0},$$

- Multi-soliton configurations can be constructed using the linear system via the "dressing" method.
- We will study the quadratic fluctuations around this solution. Invoking the semi-classical reasoning, the energy spectrum for the kink particle should be given by

$$E_{kink-sector} = 16\alpha + \frac{1}{2}\sum_{n}(\omega_n + \nu_n) + O(\alpha^2)$$

where ω_n and ν_n are the frequencies for the normal modes.

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Background Field Method				
We split the fi	elds g_+,g	by setting		

$$g_+ = g_0 e^{-i(\eta+\xi)}\,, \qquad g_- = e^{i(\eta-\xi)} g_0^{-1}\,,$$

• η, ξ are fluctuations in the static background g_0 .

• We expand $S[g_+, g_-]$ up to cubic order in η and ξ .

 $S[g_+,g_-]=S[g_0]-\int dt dy \, (\partial_\mu\eta)^2+(\partial_\mu\xi)^2+ ext{interaction terms}$

- First, we find the field equations for η and ξ and expand them to second order in θ.
- 2 Next, we expand the fluctuations in modes by assuming

$$\eta(t, \mathbf{y}) = \sum_{n} e^{i\omega_{n}t} \psi_{n}(\mathbf{y}), \quad \xi(t, \mathbf{y}) = \sum_{n} e^{i\nu_{n}t} \chi_{n}(\mathbf{y}).$$

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Equations for Fluctuations

Eigenmodes fulfill the Schrödinger-type equations:

Equations

$$\begin{bmatrix} -\partial_z^2 + V_0(z) + \theta V_1(z) + \theta^2 V_2(z) \end{bmatrix} \tilde{\psi}_n(z) = \frac{\omega_n^2}{4\alpha^2} \tilde{\psi}_n(z) ,$$
$$\begin{bmatrix} -\partial_z^2 + \theta W_1(z) + \theta^2 W_2(z) \end{bmatrix} \tilde{\chi}_n(z) = \frac{\nu_n^2}{4\alpha^2} \tilde{\chi}_n(z) .$$

With $z := 2\alpha y$, $\tilde{\psi}_n := e^{\frac{i}{4}\omega_n \theta \partial_y \varphi_0} \psi_n$, $\tilde{\chi}_n := e^{\frac{i}{4}\nu_n \theta \partial_y \varphi_0} \chi_n$ and,

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Equations for Fluctuations

Eigenmodes fulfill the Schrödinger-type equations:

Equations

$$\left[-\partial_z^2+V_0(z)+\theta V_1(z)+\theta^2 V_2(z)\right]\tilde{\psi}_n(z)=\frac{\omega_n^2}{4\alpha^2}\tilde{\psi}_n(z)\,,$$

$$\left[-\partial_z^2+\theta W_1(z)+\theta^2 W_2(z)\right]\tilde{\chi}_n(z)=\frac{\nu_n^2}{4\alpha^2}\tilde{\chi}_n(z).$$

With
$$z := 2\alpha y$$
, $\tilde{\psi}_n := e^{\frac{i}{4}\omega_n \theta \partial_y \varphi_0} \psi_n$, $\tilde{\chi}_n := e^{\frac{i}{4}\nu_n \theta \partial_y \varphi_0} \chi_n$ and,

Potentials

$$V_0 = (2 \tanh^2 z - 1), V_1 = -\omega_n^2 \frac{\sinh z}{\cosh^2 z}$$
$$V_2 = -\omega_n^2 \alpha^2 \Big(\frac{2}{\cosh^4 z} - \frac{\sinh^2 z}{\cosh^4 z}\Big)$$

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Equations for Fluctuations

Eigenmodes fulfill the Schrödinger-type equations:

Equations

$$\left[-\partial_z^2+V_0(z)+\theta V_1(z)+\theta^2 V_2(z)\right]\tilde{\psi}_n(z)=\frac{\omega_n^2}{4\alpha^2}\tilde{\psi}_n(z)\,,$$

$$\left[-\partial_z^2+\theta W_1(z)+\theta^2 W_2(z)\right]\tilde{\chi}_n(z)=\frac{\nu_n^2}{4\alpha^2}\tilde{\chi}_n(z).$$

With
$$z := 2\alpha y$$
, $\tilde{\psi}_n := e^{\frac{i}{4}\omega_n \theta \partial_y \varphi_0} \psi_n$, $\tilde{\chi}_n := e^{\frac{i}{4}\nu_n \theta \partial_y \varphi_0} \chi_n$ and,

Potentials

Potentials

$$\frac{Z}{Z}$$

$$W_1(z) = -\nu_n^2 \frac{\sinh z}{\cosh^2 z}$$

$$W_2(z) = \nu_n^2 \alpha^2 \frac{\sinh^2 z}{\cosh^4 z}$$

$$V_0 = (2 \tanh^2 z - 1), V_1 = -\omega_n^2 \frac{\sinh z}{\cosh^2 z}$$
$$V_2 = -\omega_n^2 \alpha^2 \left(\frac{2}{\cosh^4 z} - \frac{\sinh^2 z}{\cosh^4 z}\right)$$



• We have the equation

$$\left[-\partial_z^2+2\tanh^2 z-1\right]\psi_n(z)=\frac{\omega_n^2}{4\alpha^2}\psi_n(z)\,,$$

The solution consists of the discrete zero-mode

$$\psi_0(z)=\partial_z arphi_0=-rac{2}{\cosh z}\,,\quad \omega_0=0\,,$$

followed by the continuum states

 $\psi_q(z) = e^{iqz}(\tanh z - iq), \quad {}_0\omega_q^2 = 4\alpha^2(q^2 + 1), \quad q \ge 0.$

• $\psi_q(z)$ can be normalized, by putting the system in a box of length *L*.

Sportrum of Eluctuations				
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We consider θ -dependent potentials as perturbations

- $\psi_0(z) = -\frac{2}{\cosh z}$ is static, and remains a zero-mode to all order in θ .
- 2 For corrections to the spectrum of ω_n^2 we can write

$$\omega_n^2 - {}_0\omega_n^2 =: \Sigma_k \theta^k \left(\Delta_n^k(V_1) + \Delta_n^k(V_2) \right)$$

At order θ:

We observe that

$$W_1 = -\omega_n^2 rac{\sinh z}{\cosh^2 z}, \quad W_1 = -\nu_n^2 rac{\sinh z}{\cosh^2 z}$$

odd under parity. So $\Delta_n^1(V_1)$ and $\Delta_n^1(W_1)$ both vanish.

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Spectrum of Fluctuations				
At order θ^2 :				

- We find that $\Delta_n^1(V_2) \approx \frac{1}{L}$ and $\Delta_n^1(W_2) \approx \frac{1}{L}$, thus they too vanish as $L \to \infty$.
- It seems not possible to compute $\Delta_n^2(V_1)$ and $\Delta_n^2(W_1)$ analytically, but it is unlikely that they change the spectrum considerably.



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Spectrum of Fluctuations				

In the vacuum sector we have

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$$g_0 = e^{-rac{i}{2}\varphi_0} = 1$$
, $\varphi_0 = 0$, $\rho_0 = 0$,

Fluctuation equations are

$$-\partial_{\mu}\partial^{\mu}\eta - 4\alpha^{2}\eta = \mathbf{0}\,, \quad -\partial_{\mu}\partial^{\mu}\xi = \mathbf{0}\,.$$

• Thus η and ξ are plane waves:

$$\eta(t, \mathbf{y}) = \mathbf{e}^{\pm i \mathbf{k} \mathbf{y} + i \omega t}, \quad \xi(t, \mathbf{y}) = \mathbf{e}^{\pm i \mathbf{r} \mathbf{y} + i \nu t},$$

The dispersion relations

$$\omega^2 = k^2 + 4\alpha^2, \quad \nu^2 = r^2.$$

 These are in agreement with the ordinary sine-Gordon theory results.

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Spectrum of Fluctuations				

- Our perturbative treatment of noncommutativity for the spectrum of fluctuations around the kink implies no changes in the spectrum of fluctuations.
- Thus E_{kink} E_{vacuum} is in agreement with the results of the ordinary sine-Gordon model.

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Feynman Rules and Two-point functions

We move on to discuss the properties of the two-point functions of the model.

• The propagators are

$$-\!\!-\!\!-\!\!= \langle \varphi \varphi \rangle = \frac{2}{k^2 + 4\alpha^2}, \qquad = \langle \rho \rho \rangle = \frac{2}{k^2}$$

 For our purposes we only need the interactions to quadratic order in the fields ψ and ρ. The vertices are then



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Conclusions and Outlook

Feynman Rules and Two-point functions

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Feynman rules for these vertices read

$$= -\frac{1}{2^2} (k_1 \wedge k_2) \sin \left(\theta \frac{k_1 \wedge k_2}{2}\right) e^{-\frac{i}{2}\theta (k_1 \wedge k_2 + k_2 \wedge k_3)}$$
$$= \frac{1}{12} \alpha^2 e^{\left(-\frac{i}{2}\theta \sum_{i< j}^n k_i \wedge k_j\right)} - \frac{i}{2^2 \cdot 4!} k_1 \cdot (k_3 - k_2)$$
$$\times \sin \left(\theta \frac{k_2 \wedge k_3}{2}\right) e^{-\frac{i}{2}\theta (k_1 \wedge k_2 + k_1 \wedge k_3 + k_1 \wedge k_4 + k_2 \wedge k_4 + k_3 \wedge k_4)}$$

Conclusions and Outlook

Feynman Rules and Two-point functions

Feynman rules for these vertices read

•
=
$$-\frac{1}{2^2}(k_1 \wedge k_2) \sin\left(\theta \frac{k_1 \wedge k_2}{2}\right) e^{-\frac{i}{2}\theta(k_1 \wedge k_2 + k_2 \wedge k_3)}$$

•
= $\frac{1}{12} \alpha^2 e^{\left(-\frac{i}{2}\theta \sum_{i
 $\times \sin\left(\theta \frac{k_2 \wedge k_3}{2}\right) e^{-\frac{i}{2}\theta(k_1 \wedge k_2 + k_1 \wedge k_3 + k_1 \wedge k_4 + k_2 \wedge k_4 + k_3 \wedge k_4)}$$

•
$$a \wedge b = a_t b_y - a_y b_t$$



It was shown by Lechtenfeld et. al.*Nucl.Phys.B***705**(2005)) that this model do not exhibit any acausal behaviour at tree level.



- All other amplitudes, $A_{\rho\rho\to\rho\rho}$, $A_{\varphi\rho\to\varphi\rho}$, $A_{\varphi\varphi\to\rho\rho}$ and $A_{\rho\rho\to\varphi\varphi}$ vanish.
- Thus the model has no acausal effects.

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 Amplitudes for φφ → φφφφ and φφφ → φφφ also vanish. This is in agreement with the commutative sine-Gordon model.

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One-Loop Behavior				

One-loop two-point functions in vacuum sector...

• Two-point function for φ is $I_{\varphi}(P^2)$

$$I_{\varphi}(P^2) =$$
 $-$ $+$ $-$ $+$ $-$

 Non-planar diagram I₂ leads to UV/IR mixing. We observe this from

$$I_2 = \frac{-\alpha^2}{6\pi} \log \left[\alpha^2 \theta^2 P^2 + \frac{4\alpha^2}{\Lambda^2} \right] + \text{subleading terms} \,,$$

• I_3 and I_4 vanish as $\theta \to 0$. There is no UV/IR mixing due to I_1 , I_3 and I_4 .

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One-Loop Behavior

Renormalization; mass and field strength counter terms

• For $P \neq 0$, $\theta \neq 0$, the leading terms for $I_{\varphi}(P^2)$ reads

$$I_{\varphi}(P) \approx \left[rac{-lpha^2}{3\pi} + rac{P^2}{2^6\pi}
ight] \log rac{4lpha^2}{\Lambda^2} + ext{finite terms} + ext{subleading terms}$$

2 For P = 0, $I_{\varphi}(P^2)$ is the same as that of the ordinary sine-Gordon model, thus only mass renormalization is sufficient to render the theory finite.

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One-Loop Behavior				

Renormalization in the Euclidean Signature.

• We can write renormalized self-energy as:

$$\Sigma_R(P^2) = (1 + \delta Z_{\varphi})^{-1} I_{\varphi}(P^2) + \delta m_{\varphi}^2 + \delta Z_{\varphi} P^2$$

and assume the renormalization conditions

$$\Sigma_R(P^2)\Big|_{P^2=P_0^2}=0, \quad \frac{d}{dP^2}\Sigma_R(P^2)\Big|_{P^2=P_0^2}=0$$

From these considerations we find for δm²_φ and δZ_φ:

$$\delta m_{\varphi}^{2} = \frac{1}{1 + \delta Z_{\varphi}} \left[\frac{\alpha^{2}}{3\pi} \log \frac{4\alpha^{2}}{\Lambda^{2}} \right], \quad \delta Z_{\varphi} = \frac{-1 + \sqrt{1 - \frac{1}{2^{4}\pi} \log \frac{4\alpha^{2}}{\Lambda^{2}}}}{2}$$

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One-Loop Behavior				

For ρ , the one-loop two-point function is

$$I_{
ho}(\mathcal{P}^2) = \left(rac{1}{2}I_3 + I_4
ight)\Big|_{4lpha^2
ightarrow \mu^2}$$

- μ is a small mass for ρ introduced to regularize the IR limit.
- $I_{\rho}(P^2)$ is present purely due to the noncommutativity: $I_{\rho}(P^2) \rightarrow 0$ as $\theta \rightarrow 0$. However, it does not lead to any UV/IR mixing.
- There is no mass renormalization and the field-strength renormalization is given by

$$\delta Z_{
ho} = rac{-1 + \sqrt{1 - rac{3}{2^5 \pi} \log rac{\mu^2}{\Lambda^2}}}{2}$$



- **Remark1:** We stress that, these results are valid for $\theta \neq 0$. When $\theta \to 0$, in $I_{\varphi}(P^2)$ and $I_{\rho}(P^2)$, the divergent terms in Λ cancel with those in θ . In this case, the standard answer for the commutative sine-Gordon model is recovered, and a mass counter term for the field φ is sufficient to renormalize the theory.
- **Remark2:** When *I*(*P*²) are analytically continued to the Minkowski space, the logarithms develop branch cuts. This leads to imaginary parts in the total one-loop amplitudes, and for space-like external momenta to the violation of unitarity, as the optical theorem is no longer satisfied.

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- We have studied the quantum aspects of sine-Gordon model in noncommutative spacetime. Our aim has been to infer to what extent the classical integrability is useful in this respect.
 - We have presented a perturbative treatment of noncommutativity to study the spectrum of fluctuations around the kink. This implied the latter is in good agreement with the ordinary sine-Gordon model.
- We have found that two-point functions at one-loop level show some interesting features.
 - There is UV/IR mixing due to interactions coupled with α², but it appears that there are non-planar diagrams which do not lead to UV/IR mixing effects.
 - We have exhibited the mass and field strength renormalizations in Euclidean signature. However, in Minkowski signature time-space noncommutatvity still causes unitarity violation.
 - Although, the usual vacuum subtraction can be performed it is not clear, how to regularize the divergences of the theory in Minkowski space.

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- 3 It maybe be helpful to study the quantum effects in the 2 + 1-dimensional Ward-model to gain more insights to the structure of the present class of models.
- 4 It will certainly be useful to study the SUSY generalizations of this model and see if it helps in regularizing the divergences of the bosonic theory. Investigations in this direction are already underway.

A First Guess. A NC sine-Gordon action and why it is not useful.

Consider the action obtained by deforming all products to *-products in the commutative theory.

$$S = \int dt dy \frac{1}{2} \partial_{\mu} \phi \star \partial^{\mu} \phi + 2\alpha^{2} (\cos_{\star} \phi - 1)$$

 $\cos_{\star}\phi = 1 - \frac{1}{2}\phi \star \phi + \cdots$. The field equation becomes

$$\partial_{\mu}\partial^{\mu}\phi = -4\alpha^{2}\sin_{\star}\phi.$$

This action is no good, for

 The associated currents are deformed in the same way, but they are no longer conserved!.

- Normal ordering of the interaction is not sufficient to cancel the divergences any more.
- Consider, for example, the two-point functions. They come in three different kinds: planar, non-planar and mixed.

planar :
non-planar :

$$\sum_{non-planar} \approx \log \left[\alpha^2 \theta^2 P^2 + \frac{4\alpha^2}{\Lambda^2} \right]$$

But a mixed diagram has sub-diagram(s) which are planar

$$\boxed{ \qquad } \approx \log\left[\frac{4\alpha^2}{\Lambda^2}\right] \log\left[\alpha^2\theta^2 P^2 + \frac{4\alpha^2}{\Lambda^2}\right]$$

 Thus, some diagrams get coefficients depending on the external momenta *P*, and it is not possible to sum the counter terms to get a cos_{*} φ interaction.

