
# Statistics Short Exercise(s)

Mario Pelliccioni
Luca Lista

CMS DAS – Pisa 2019

# Welcome!

A 4 hours class

Covering a few most relevant use-cases for statistical analysis in HEP

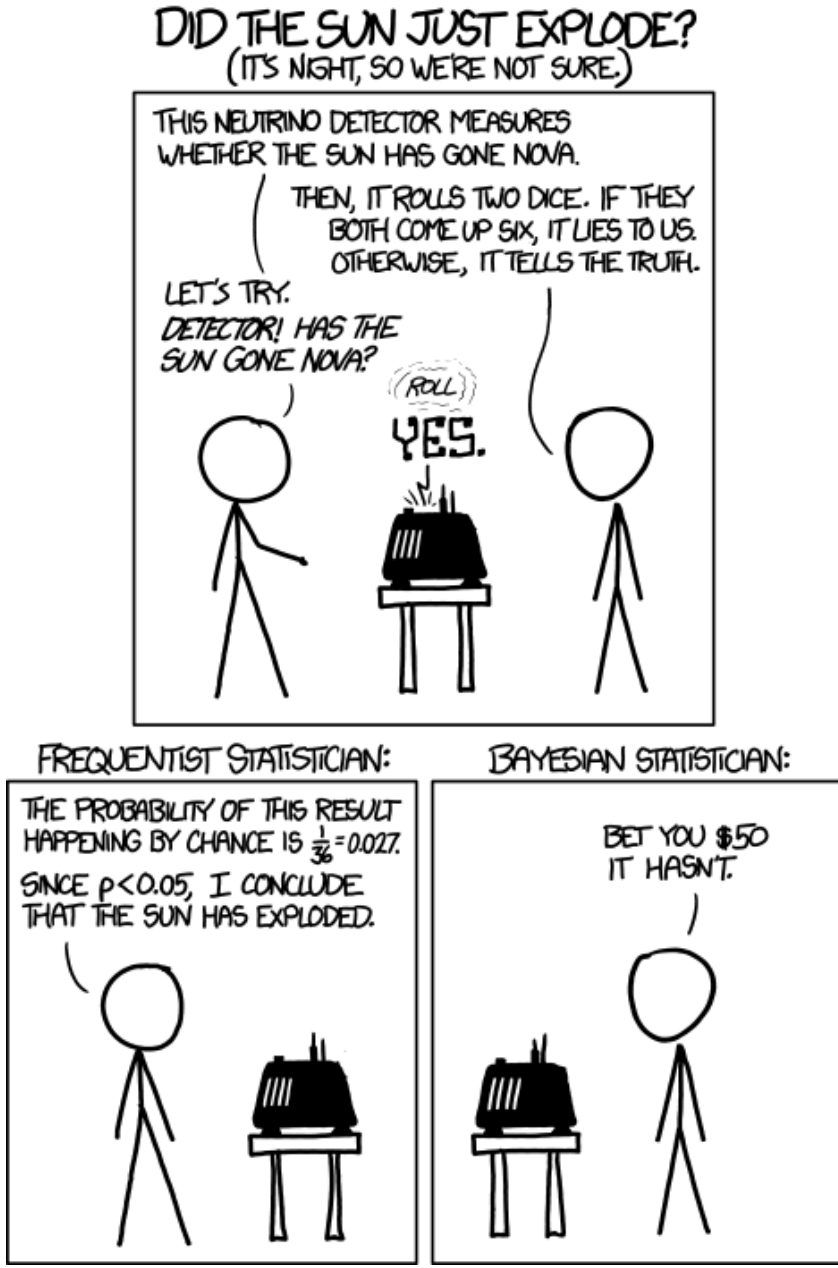
Using RooFit and RooStats as main tools

You can use your laptops for the exercises (provided you installed ROOT with the --enable-roofit option)

CERN/other labs central clusters normally work too

Exercises will be in PyROOT, so python installation necessary

We will flash a few introductory slides for each topic



# Disclaimer

The point of this class is to introduce you to some libraries that let you use different statistical tools

We will try to introduce as many different tools as we can

These are not the best (or most appropriate) ways to approach **any** statistical problem

It's your responsibility to find (or build) the best tool to do your job!

# RooFit and RooStats

**RooFit:** a ROOT library containing classes that allow to perform multi-dimensional (un)binned maximum likelihood/chi2 fits, toy-MC generation, plotting, etc

**RooStats:** a ROOT library that uses RooFit and provides classes to perform statistical interpretation of your results

# Documentation

For most of what I do, I refer to the ROOT reference guide:

https://root.cern.ch/doc/master/classes.html

This includes RooFit and RooStats reference

RooFit manual (a bit outdated):

https://root.cern.ch/download/doc/RooFit_Users_Manual_2.91-33.pdf

RooStats documentation

https://twiki.cern.ch/twiki/bin/view/RooStats/WebHome

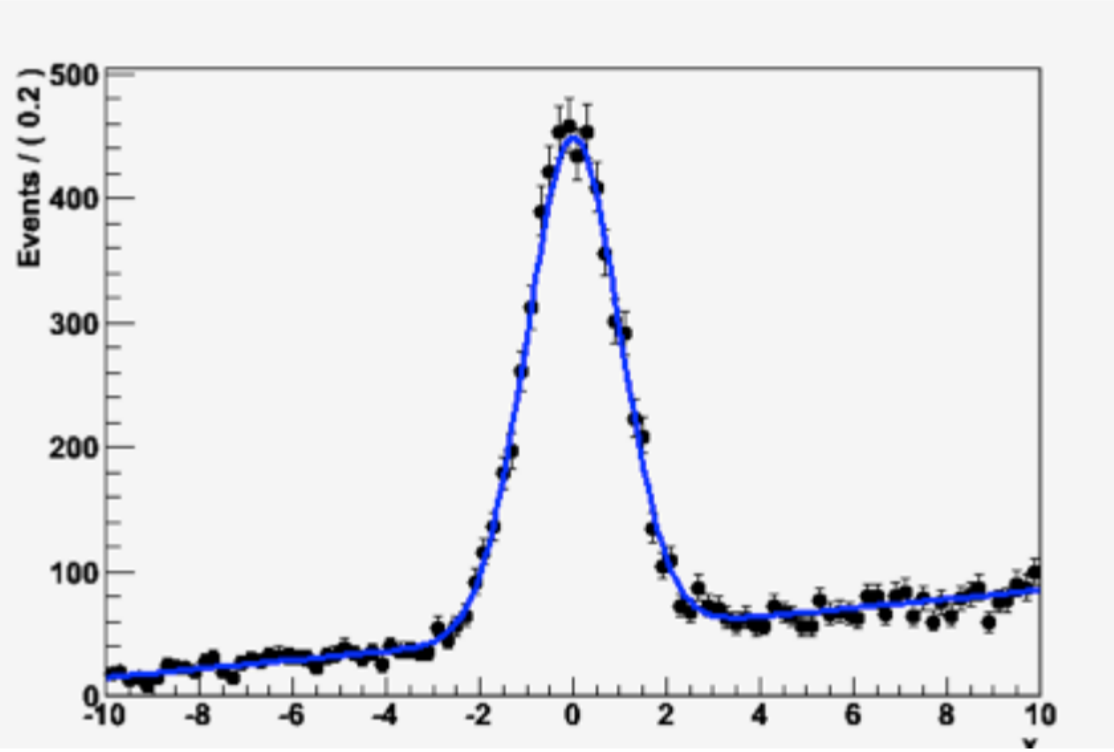
More RooFit/RooStats examples

https://github.com/pellicci/UserCode/tree/master/RooFitStat_class (C++ based)

https://github.com/pellicci/UserCode/tree/master/RooFitStat_class_python

# Why do we need RooFit?

- Focus on one practical aspect of many data analysis in HEP: How do you formulate your p.d.f. in ROOT
  - For 'simple' problems (gauss, polynomial) this is easy




  - But if you want to do unbinned ML fits, use non-trivial functions, or work with multidimensional functions you quickly find that you need some tools to help you

# The origins

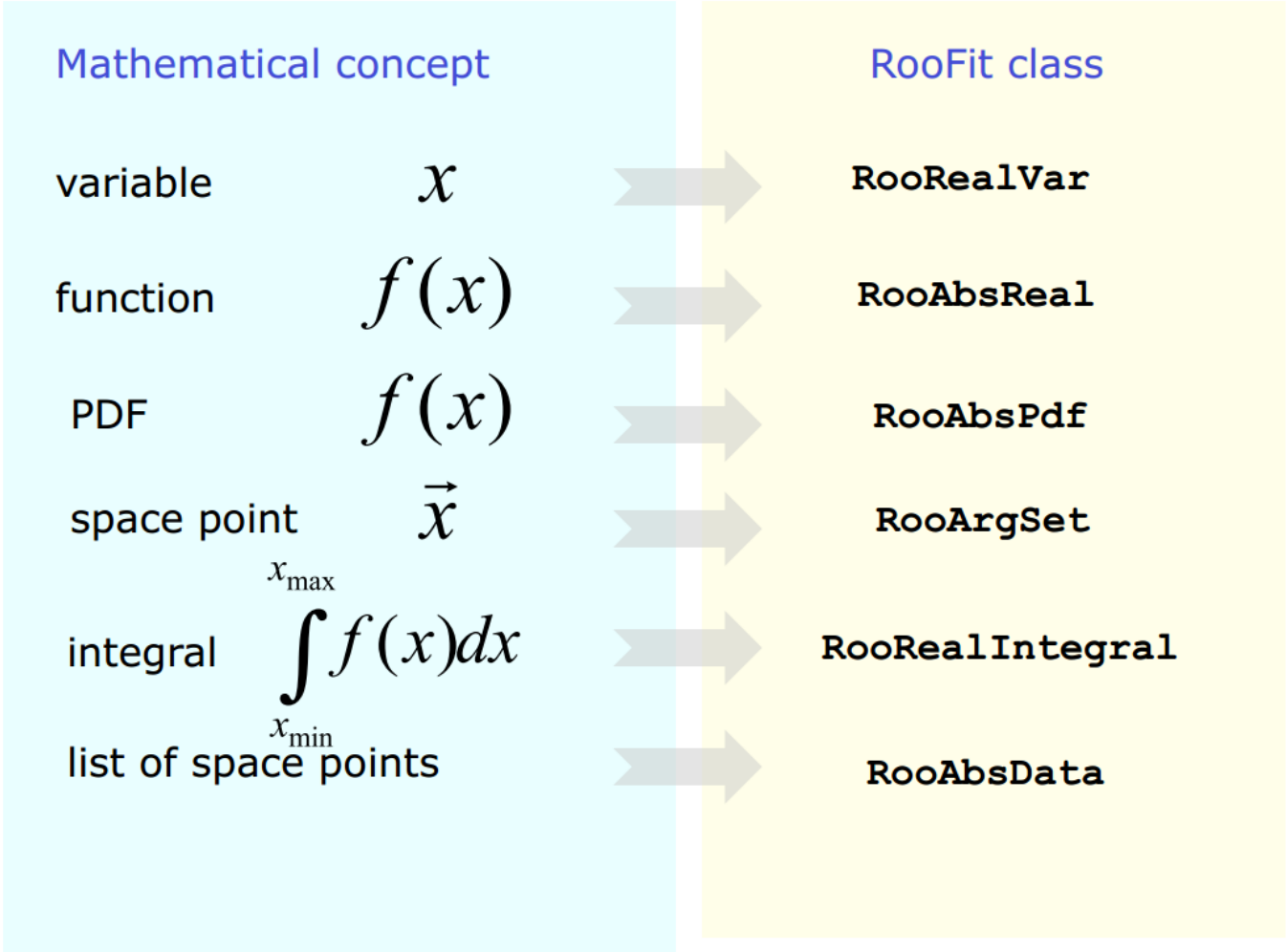
- BaBar experiment at SLAC: Extract sin(2β) from time_ dependent CP violation of B decay: $e^+e^-$ → Y(4s) → BB
  - Reconstruct both Bs, measure decay time difference
  - Physics of interest is in decay time dependent oscillation

$$f_{sig} \cdot \left[ \mathrm{SigSel}(m; \vec{p}_{sig}) \cdot \left( \mathrm{SigDecay}(t; q_{sig}, \sin(2\beta)) \otimes \mathrm{SigResol}(t \mid dt; r_{sig}) \right) \right] +$$
$$(1 - f_{sig}) \left[ \mathrm{BkgSel}(m; \vec{p}_{bkg}) \cdot \left( \mathrm{BkgDecay}(t; q_{bkg}) \otimes \mathrm{BkgResol}(t \mid dt; r_{bkg}) \right) \right]$$

- Many issues arise
  - Standard ROOT function framework clearly insufficient to handle such complicated functions → must develop new framework
  - Normalization of p.d.f. not always trivial to calculate → may need numeric integration techniques
  - Unbinned fit, >2 dimensions, many events → computation performance important → must try optimize code for acceptable performance
  - Simultaneous fit to control samples to account for detector performance

# “Dictionary”

- Mathematical objects are represented as C++ objects




RooFit uses MINUIT for most of its work, it just provides an easy to use interface and optimizations

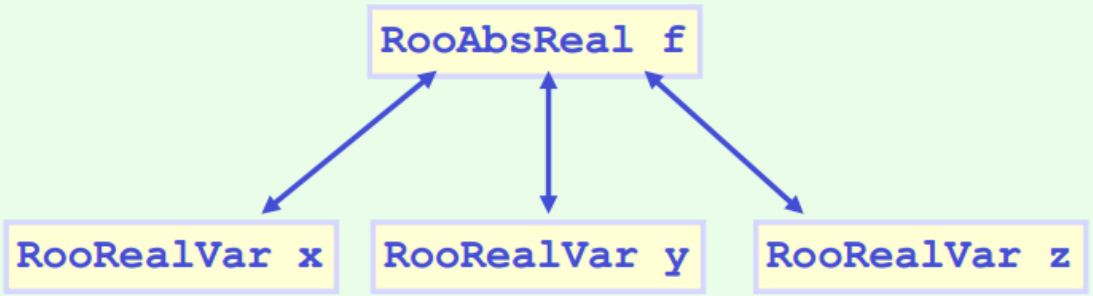
# Design philosophy

- Represent relations between variables and functions as client/server links between objects

Math

$$f(x,y,z)$$

RooFit diagram




RooFit code

```
RooRealVar x("x","x",5) ;
RooRealVar y("y","y",5) ;
RooRealVar z("z","z",5) ;
RooBogusFunction f("f","f",x,y,z) ;
```

# Variables

All variables (observables or parameters) are defined as **RooRealVar**

Several constructors available, depending on the needs:

```
var1 = ROOT.RooRealVar("var1","My first var",4.15)        //constant variable
var2 = ROOT.RooRealVar("var2""My second var",1.,10.); //range, no initial value
var3 = ROOT.RooRealVar("var3""My third var",3.,1.,10.); //valid range, initial value
```

You can also specify the unit (mostly for plotting purposes)

```
time = ROOT.RooRealVar("time","Decay time",0.,100.,"[ps]");
```

You can change the properties of your RooRealVar later (setRange, setBins, etc.)

If you want to be 100% sure a variable will stay constant, use RooConstVar

# Probability Density Functions

Each PDF in RooFit must inherit from RooAbsPdf

RooAbsPdf provides methods for numerical integration, events generation (hit & miss), fitting methods, etc.

RooFit provides a very extensive list of predefined functions (RooGaussian, RooPolynomial, RooCBShape, RooExponential, etc…)

If possible, always use a predefined function (if analytical integration or inversion method for generation are available, it will speed your computation)

You can always define a custom function using RooGenericPdf

# Data Handling

Two basic classes to handle data in RooFit:

- **RooDataSet**: an unbinned dataset (think of it as a TTree). An ntuple of data

- **RooDataHist**: a binned dataset (think of it as a THXF)

Both types of data handlers can have multiple dimensions, contain discrete variables, weights, etc.

# The perfect container

In order to “move” information among different RooFit/RooStats programs, one can use the RooWorkspace class

A **RooWorkspace** can contain:

- Variables
- PDFs
- DataSets

A RooWorkspace can be saved into a ROOT file

We’ll see how to use it.

# The problem at hand

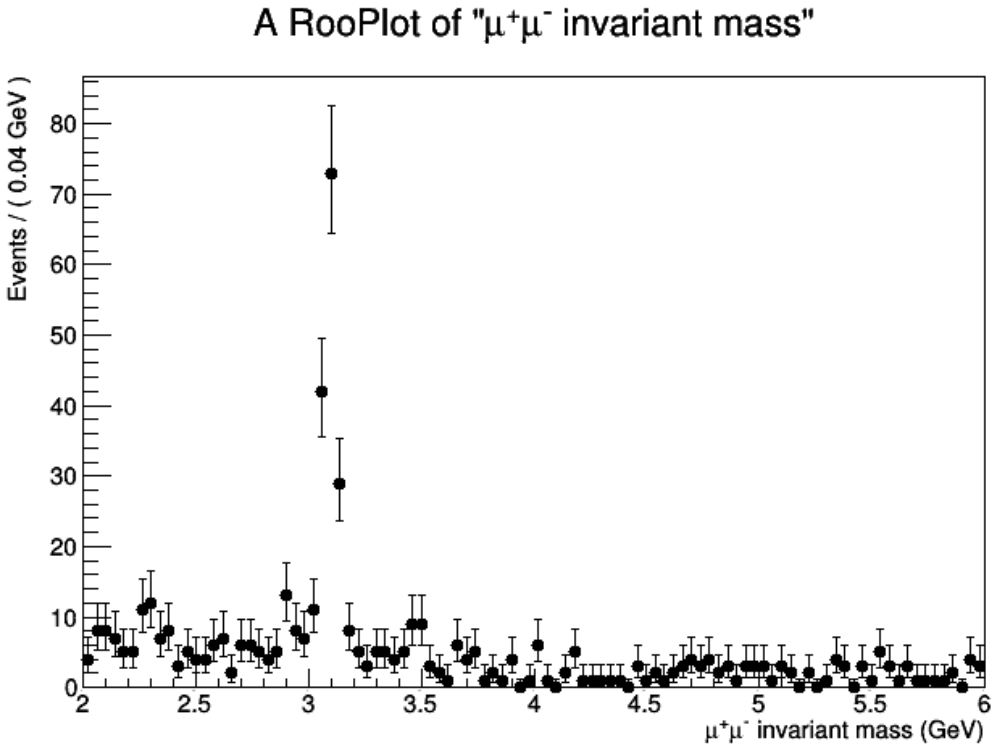
We will be analyzing a sample from the 2010 CMS data taking

All CMS data from Run1 is public → opendata.cern.ch

- Events with two opposite sign muons
- Calculated the invariant mass of the system
- Saved it into a RooDataSet (a 1D ntuple containing “mass” variable)

First, let’s look at the first three weeks of data taking (corresponds to about half a $pb^{-1}$ of integrated lumi)

We’ll be studying this distribution



# Exercise #0

The first exercise involves RooFit only

- Construct a J/ψ and ψ(2S) + background PDF
  - J/ψ with a Crystal Ball function
  - ψ(2S) with a Gaussian function
  - Background with a polynomial
- For now, the ψ(2S) will involve a very small amount of signal events
- Fit it, plot it, save it

We are going to use this program all the way through the exercises

# Parameter of interest

A parameter of interest is a variable that you want to know to the best precision and accuracy possible. It depends on the problem

Number of $\psi$(2S) could be considered the POI of the problem

In reality, we'd probably be more interested in cross section of $\psi$(2S) production → real connection with theory

$$\sigma(pp \rightarrow \psi(2S))) * BR(\psi(2S) \rightarrow \mu\mu) = \frac{N_{\psi(2S)}}{\mathcal{L} * \epsilon_{\mu\mu}}$$

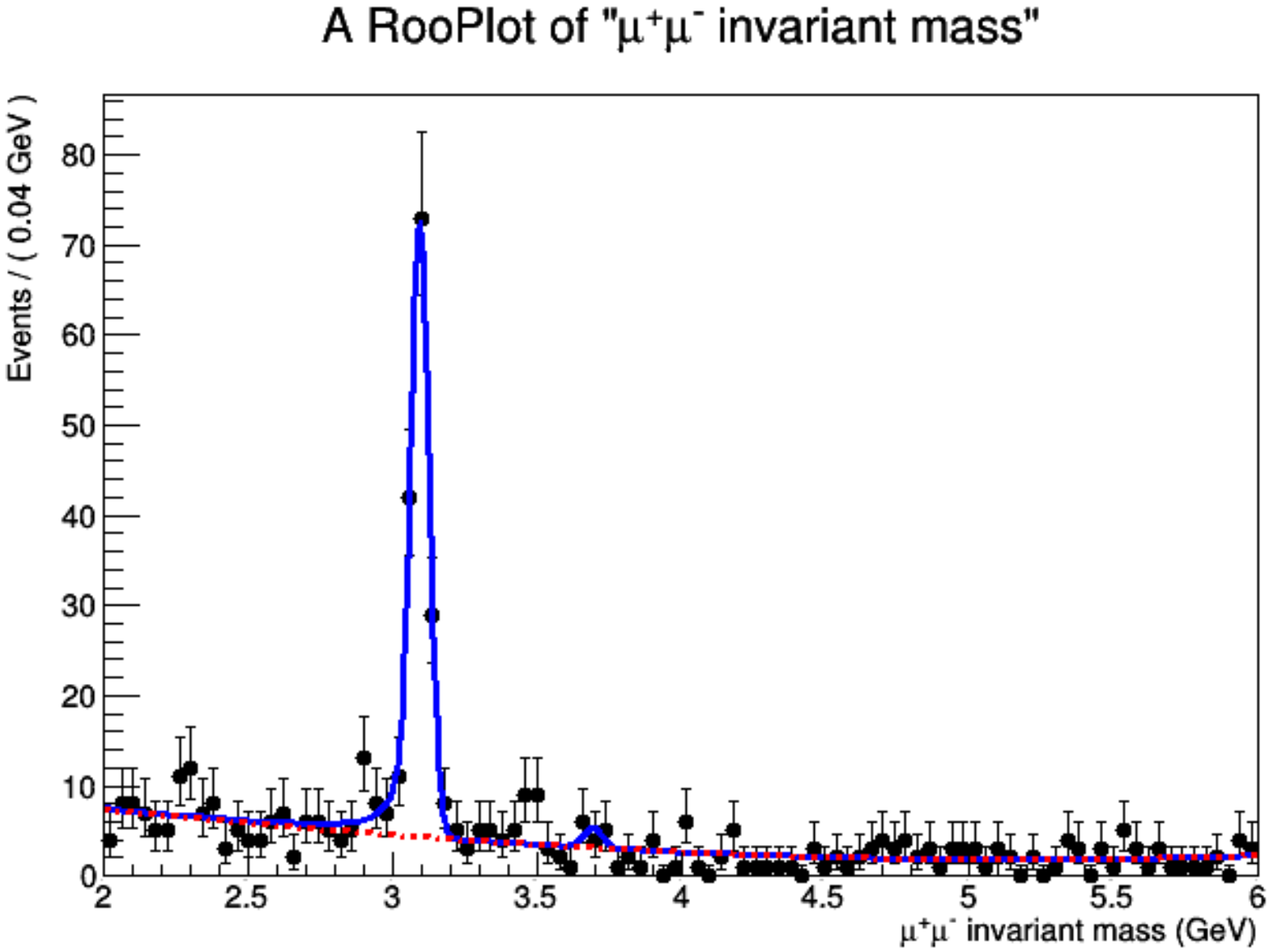
How do we express our problem in this way?

We'll assume:

75% total efficiency
A luminosity of 0.64 $pb^{-1}$

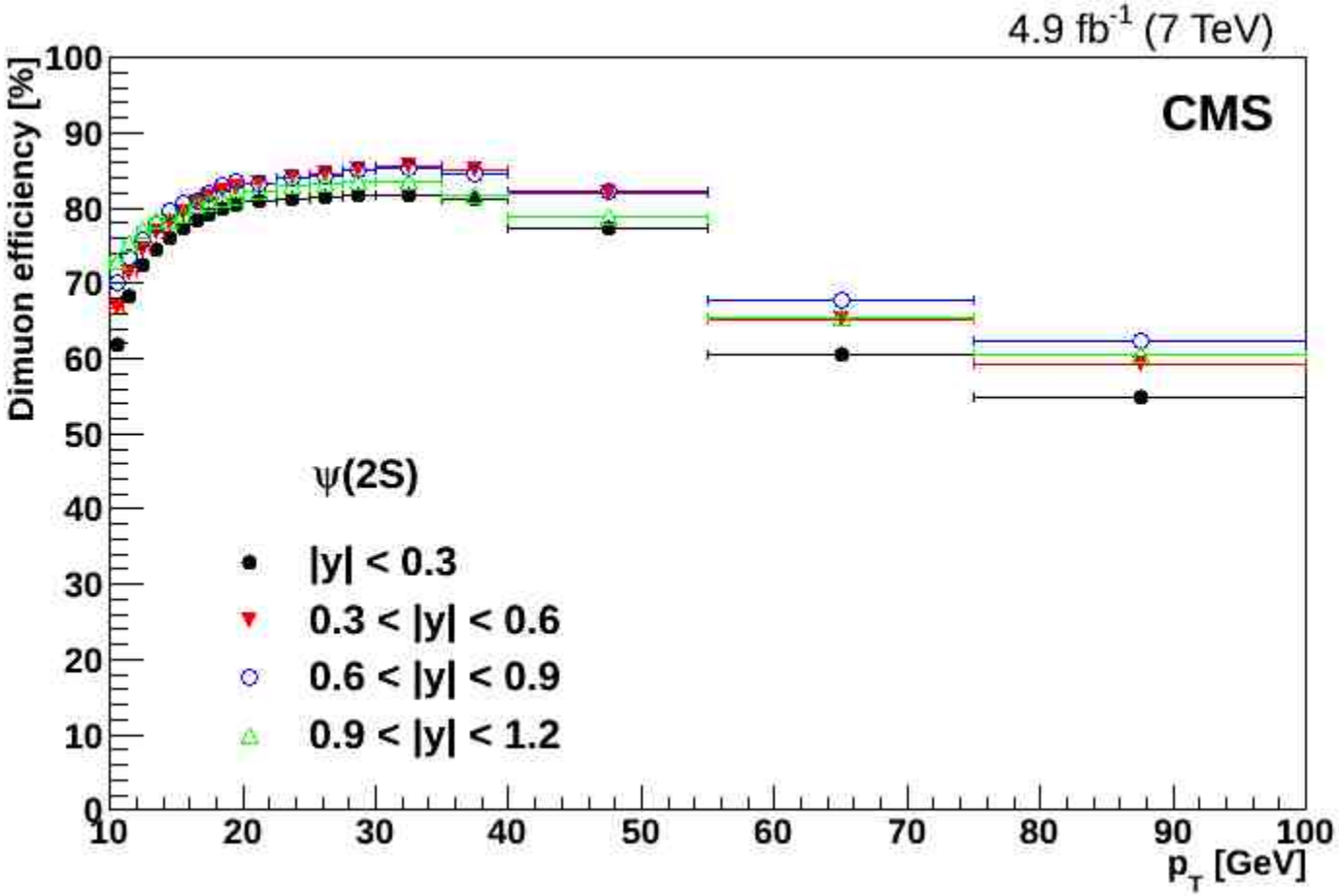
Both efficiency and luminosity uncertainties are negligible

# Result of exercise #0



# Dimuon efficiency

From CMS-BPH-14-001

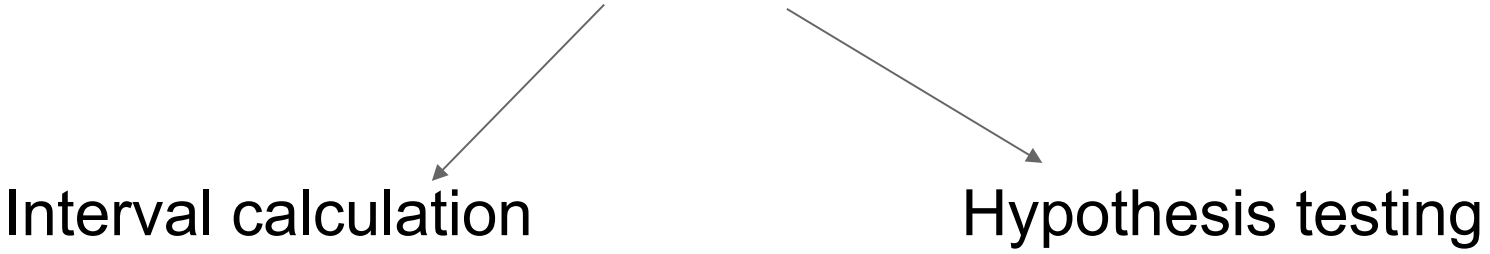


# RooStats

Set of libraries for statistical interpretation of your results

→ communicates with RooFit via RooWorkspace

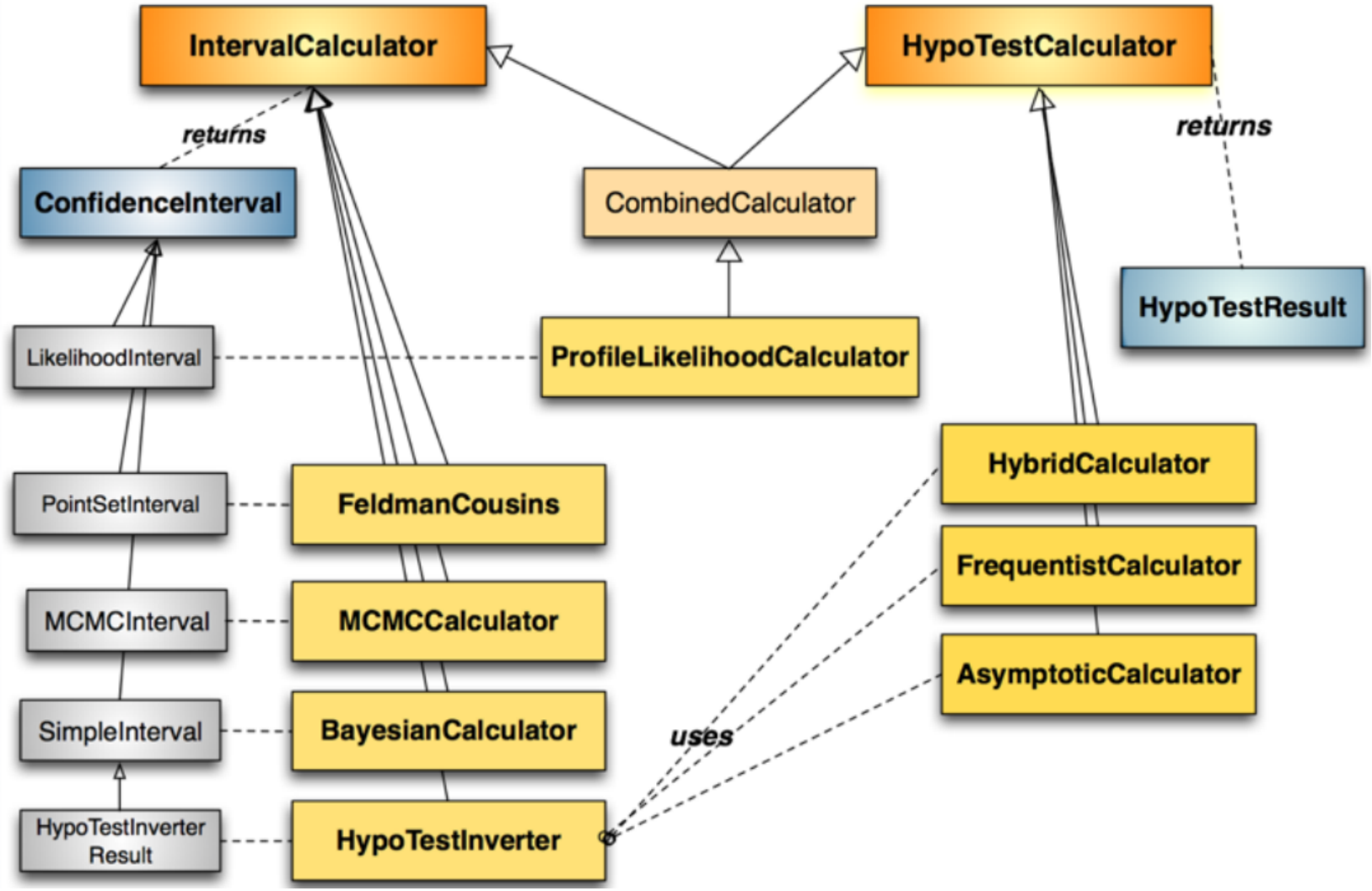
RooStats does essentially two things:




To do this, it uses “calculators”

# RooStats design

C++ classes that reproduce statistical concepts



# Main RooStats Calculators

## ProfileLikelihood calculator

- interval estimation using asymptotic properties of the likelihood function

## Bayesian calculators

- interval estimation using Bayes theorem

`BayesianCalculator` (analytical or adaptive numerical integration)

`MCMCCalculator` (Markov-Chain Monte Carlo)

## HybridCalculator, FrequentistCalculator

- frequentist hypothesis test calculators using toy data (difference in treatment of nuisance parameters)

## AsymptoticCalculator

- hypothesis tests using asymptotic properties of likelihood function

## HypoTestInverter

- invert hypothesis test results (from Asympototic, Hybrid or FrequentistCalculator) to estimate an interval
- main tools used for limits at LHC (limits using CLs procedure)

## NeymanConstruction and FeldmanCousins

- frequentist interval calculators

# Exercise #1

Exercise #0 told us that there's clearly no significant peak in the distribution

Is this actually clear? How do we quantify?

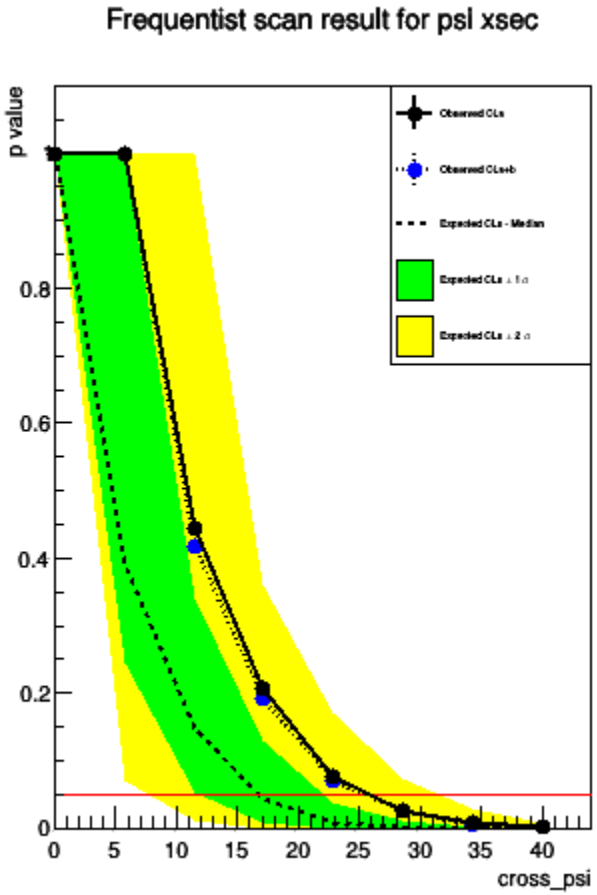
# Exercise #2

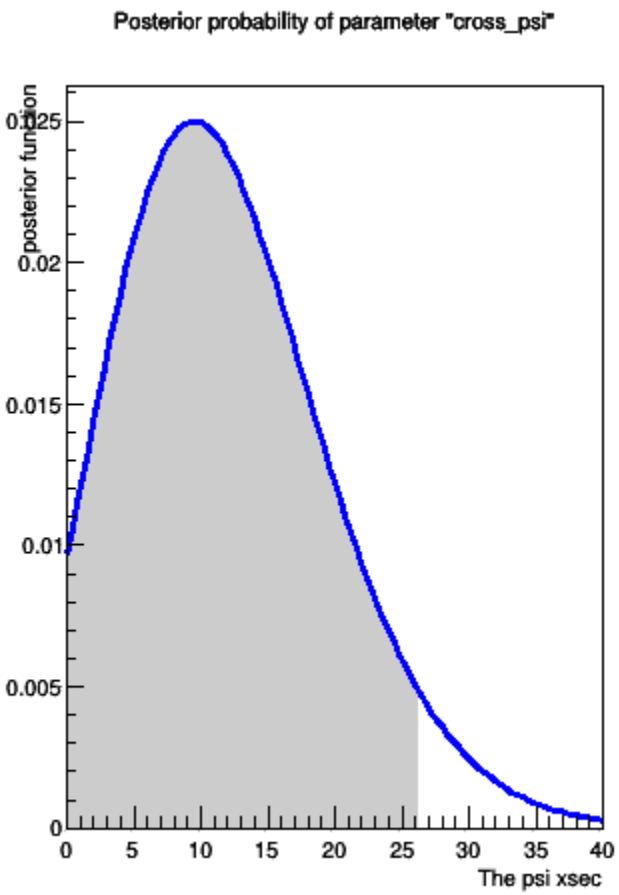
From exercise #1 we know that our excess is “not significant”.

The normal procedure here is to evaluate an upper limit on our parameter of interest.

For the frequentist method, we will use $CL_s$...

# Result of exercise #2



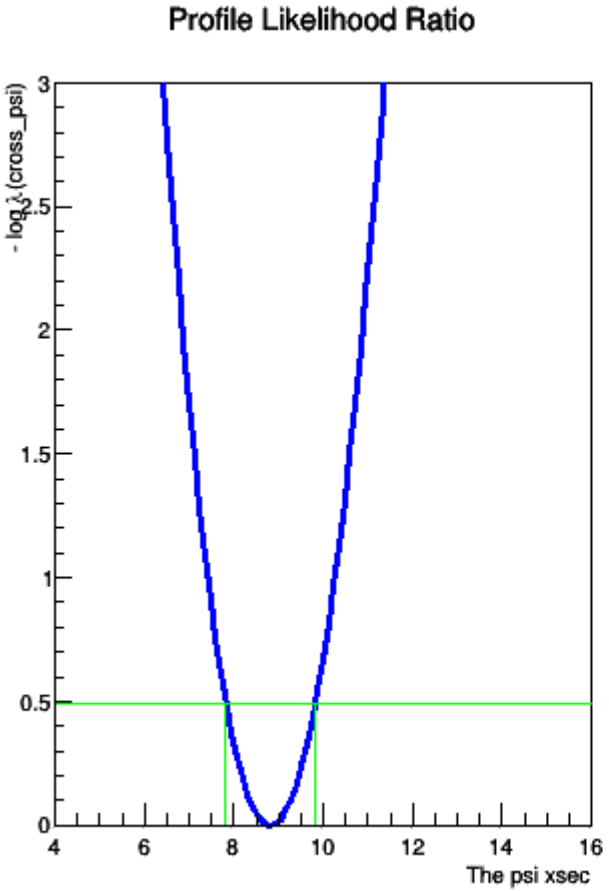
# Exercise #3

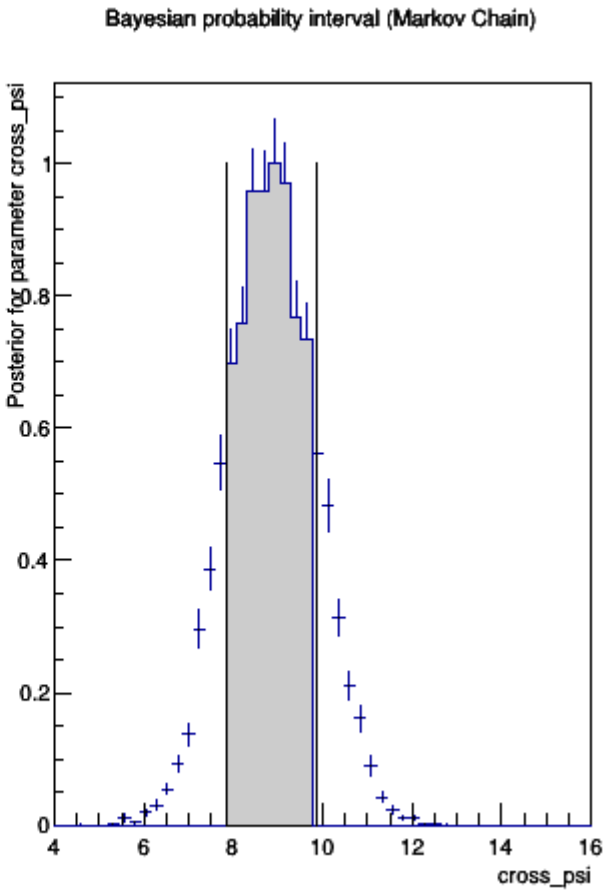
Let's now go to a scenario where we have a significant excess

- Get the full 2010 statistics file
- Rerun exercise 0 and 1 to recreate the workspace and calculate the new significance

Now we can measure the properties of our discovery

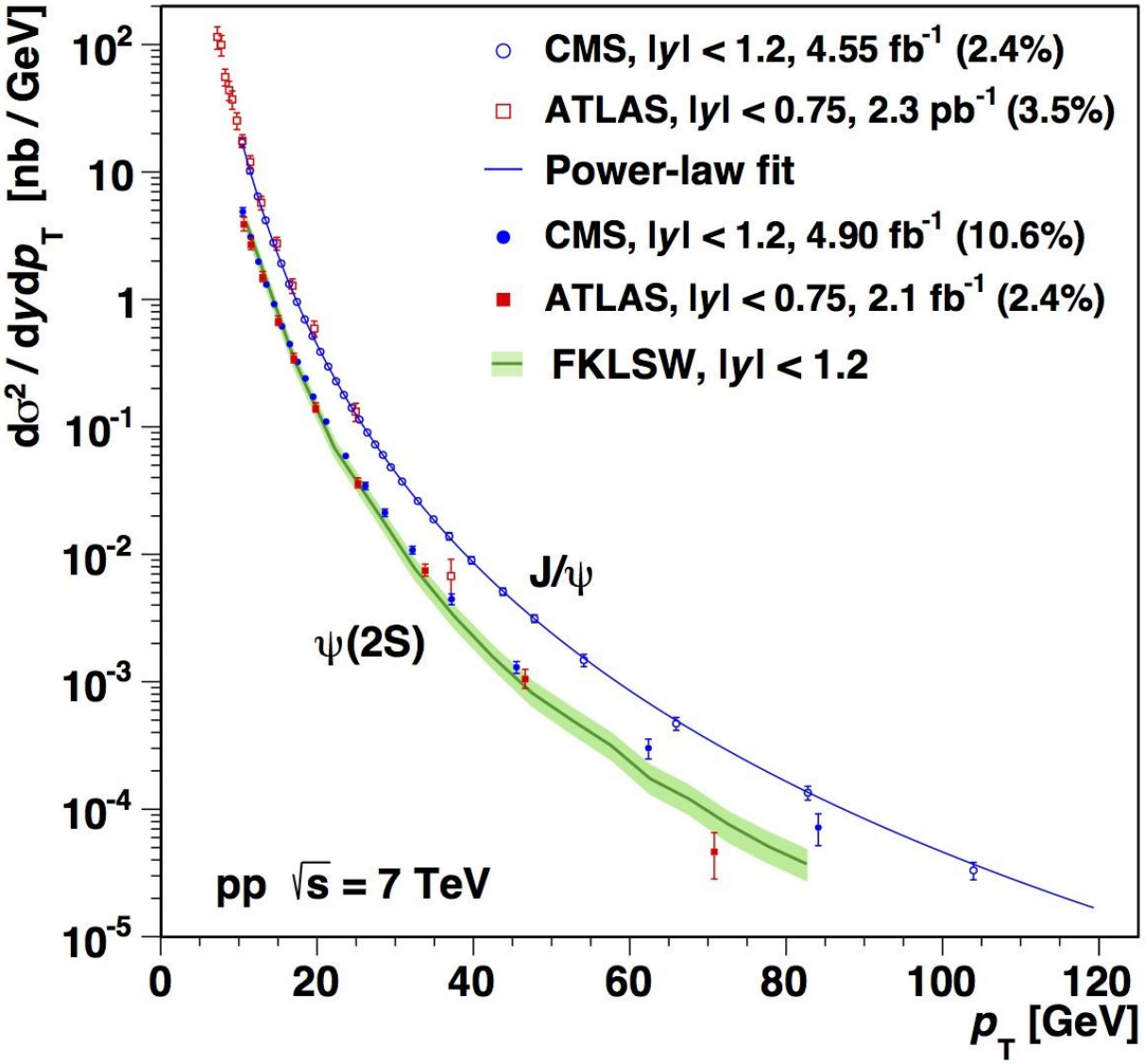
# Result of exercise #3



# ψ(2S) cross section

From CMS-BPH-14-001




Remember that

BR(Ψ(2S)→μμ ) ~ $8*10^{-3}$

# That’s all folks!